



NOvA

NuMI Off-axis ν_e Appearance Experiment

Karol Lang
University of Texas at Austin

For the NOvA Collaboration



**XII International Workshop on
“Neutrino Telescopes”**

March 6-9, 2007
"Istituto Veneto di Scienze, Lettere ed Arti", Palazzo Franchetti - Campo S. Stefano
Venice



Outline



□ NO ν A physics

- ◆ $\nu_\mu \rightarrow \nu_e$
- ◆ mass hierarchy
- ◆ CP violation
- ◆ $\nu_\mu \rightarrow \nu_\mu$

□ Experimental strategy

□ Accelerator & Beam

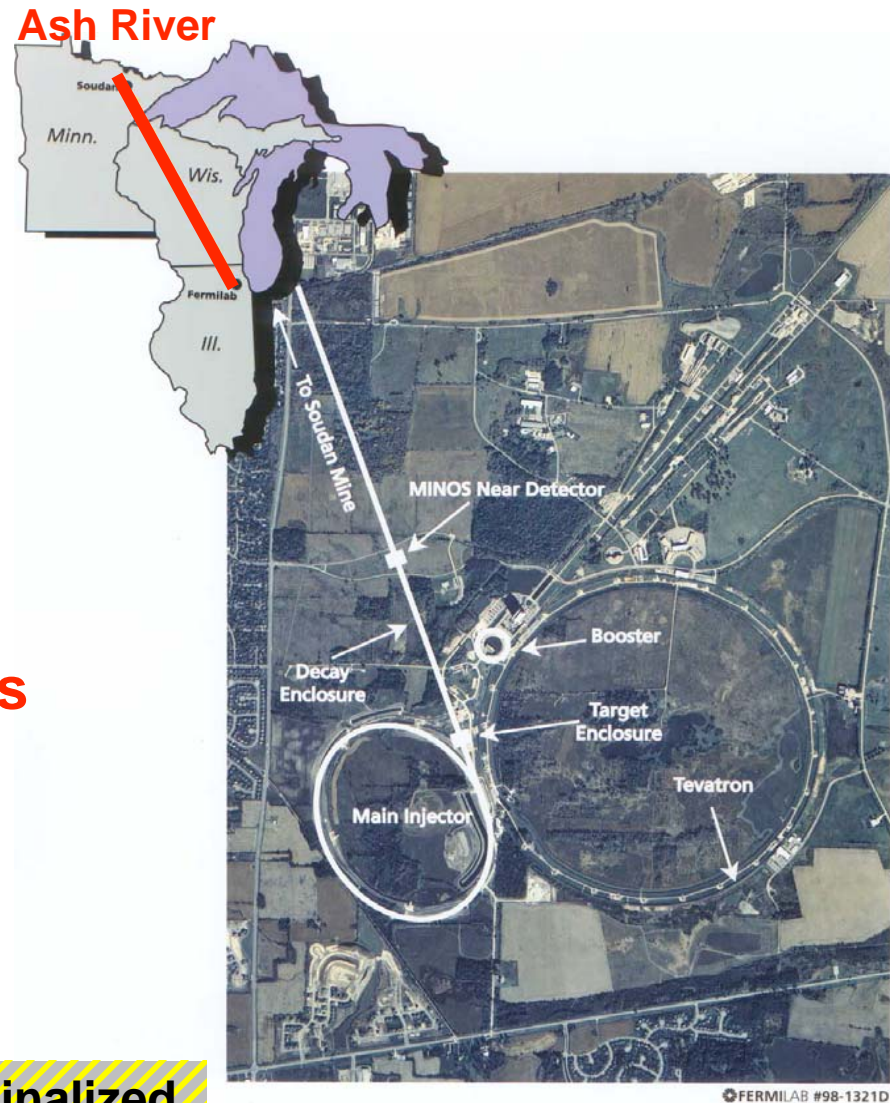
- ◆ current status
- ◆ future upgrades

□ Far and Near NO ν A detectors

- ◆ technology
- ◆ update on R&D

□ Summary & Outlook

Technical Design Report (TDR) being finalized.
All plots will be soon updated.



NuMI at
Fermilab and Minnesota



Sub-dominant mixing & CP-violating phase



Pontecorvo – Maki – Nakagawa – Sakata (PMNS) matrix

$$|\nu_\ell\rangle = \sum U_{\ell m} |\nu_m\rangle$$

$$U = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

Atmospheric

$$\nu_\mu \rightarrow \nu_\tau$$

Atmospheric

$$\nu_e \rightarrow \nu_\mu, \nu_\tau$$

Solar

$$c_{ij} = \cos\theta_{ij}$$

$$s_{ij} = \sin\theta_{ij}$$

$$U = \begin{pmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta} & s_{23}c_{13} \\ s_{12}s_{23} - c_{12}c_{23}s_{13}e^{i\delta} & -c_{12}s_{23} - s_{12}c_{23}s_{13}e^{i\delta} & c_{23}c_{13} \end{pmatrix}$$



Mass hierarchy via oscillations in matter



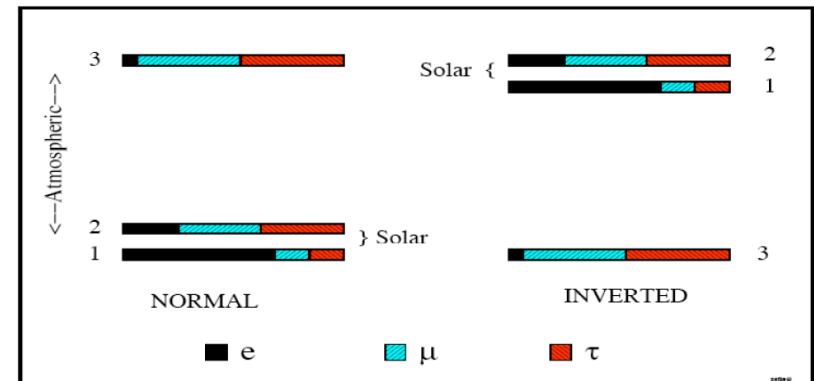
□ In matter at oscillation maximum

$$P_{matter}(\nu_\mu \rightarrow \nu_e) = \left(1 \pm 2 \frac{E}{E_R}\right) P_{vacuum}(\nu_\mu \rightarrow \nu_e)$$

□ + for neutrinos with normal mass hierarchy and anti-neutrinos with inverted mass hierarchy

□ – vice versa

□ About a +/- 30% for NuMI to Ash River



$$P_{vacuum}(\nu_\mu \rightarrow \nu_e) \approx \sin^2 \theta_{23} \sin^2 2\theta_{13} \sin^2 \Delta_{atm} = 2.5\% \left(\frac{\sin^2 2\theta_{13}}{0.05} \right)$$

$$E_R = \frac{\Delta m_{32}^2}{2\sqrt{2}G_F N_e} \approx 11 \text{ GeV}$$

$$E_{1st \max} = 1.64 \text{ GeV} \left(\frac{\Delta m_{32}^2}{2.5 \times 10^{-3} \text{ eV}^2} \right) \left(\frac{L}{810 \text{ km}} \right)$$

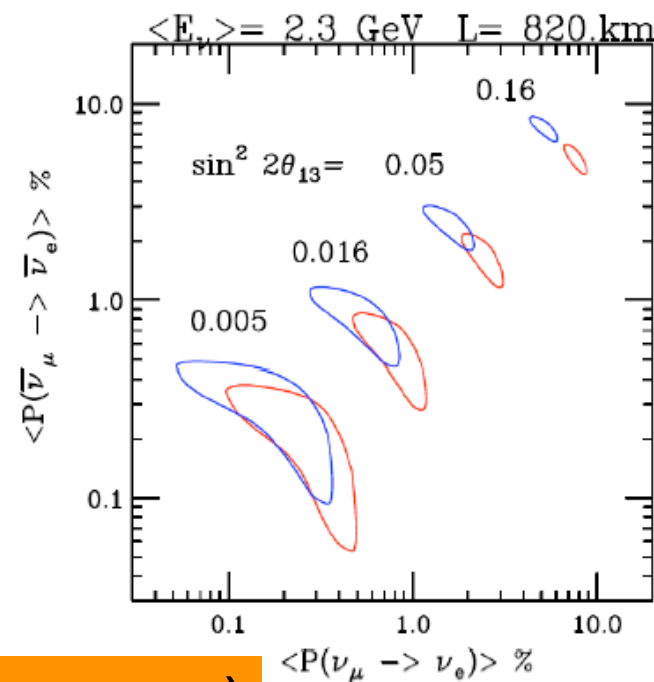
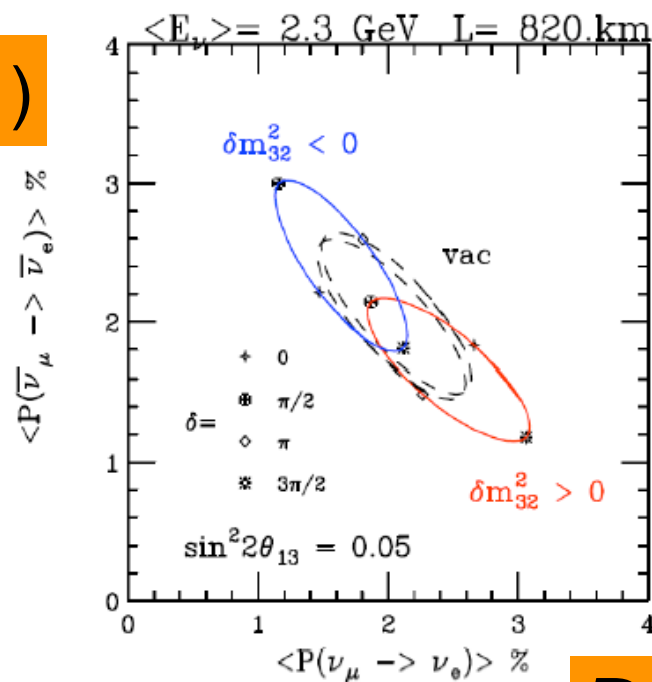
$$\Delta_{atm} \approx 1.27 \left(\frac{\Delta m_{32}^2 (\text{eV}^2) L (\text{km})}{E (\text{GeV})} \right)$$



$\sin^2 2\theta_{13}$ versus mass hierarchy

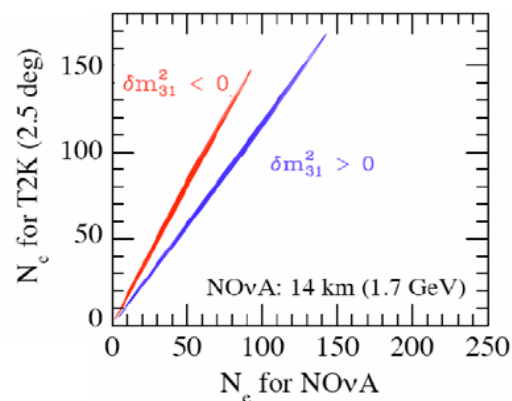


$$P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)$$



$$P(\nu_\mu \rightarrow \nu_e)$$

- ❑ Ambiguities
- ❑ Other measurements helpful (e.g., T2K)



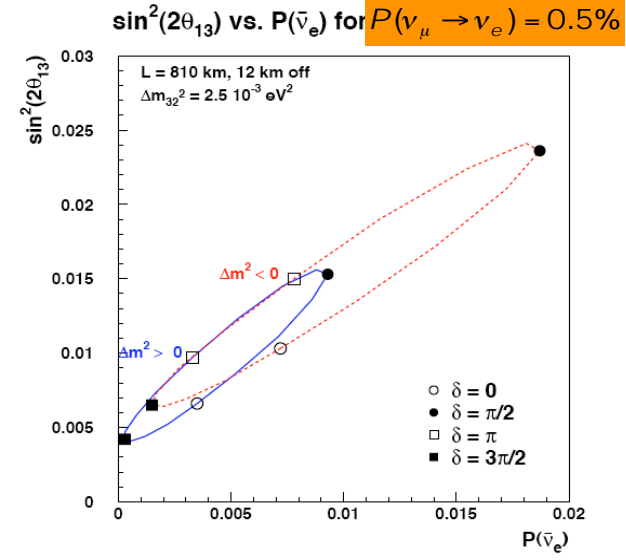
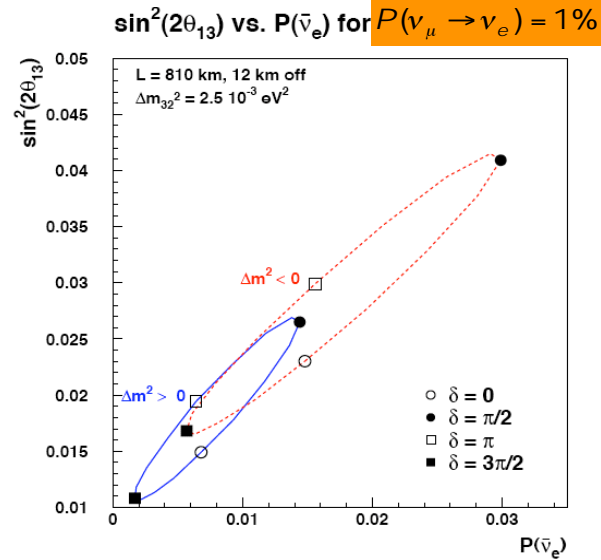
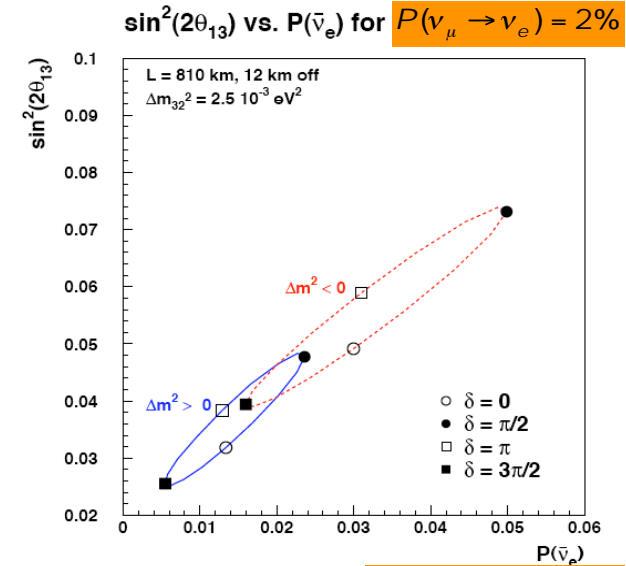
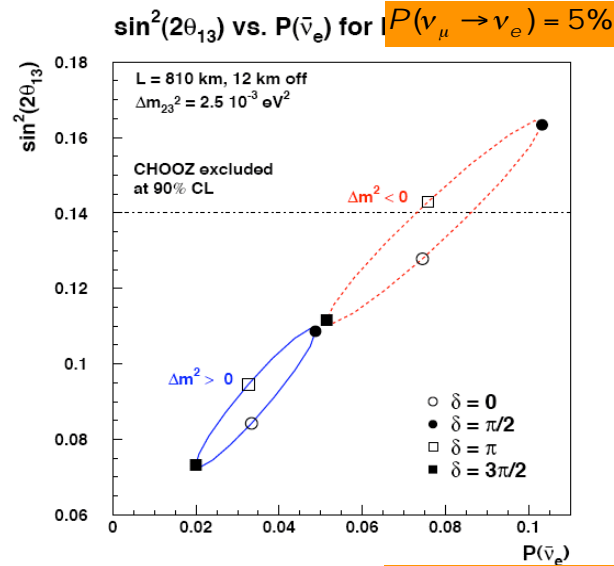


Ambiguities versus

$$P(\nu_\mu \rightarrow \nu_e)$$



$$\sin^2(2\theta_{13})$$



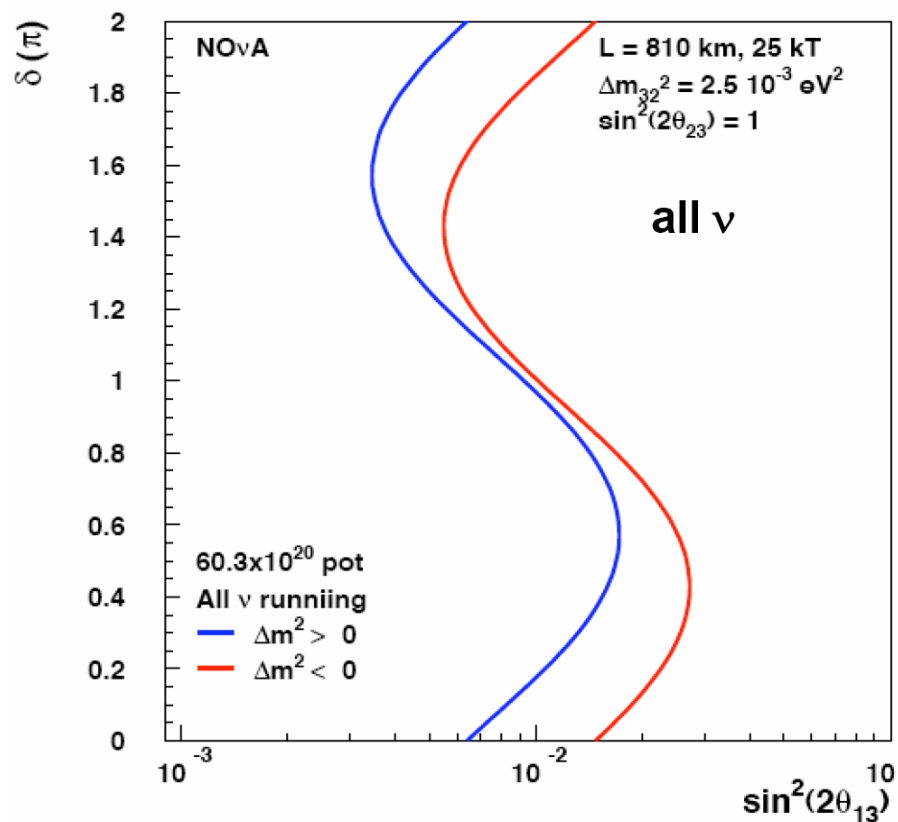
$$P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)$$



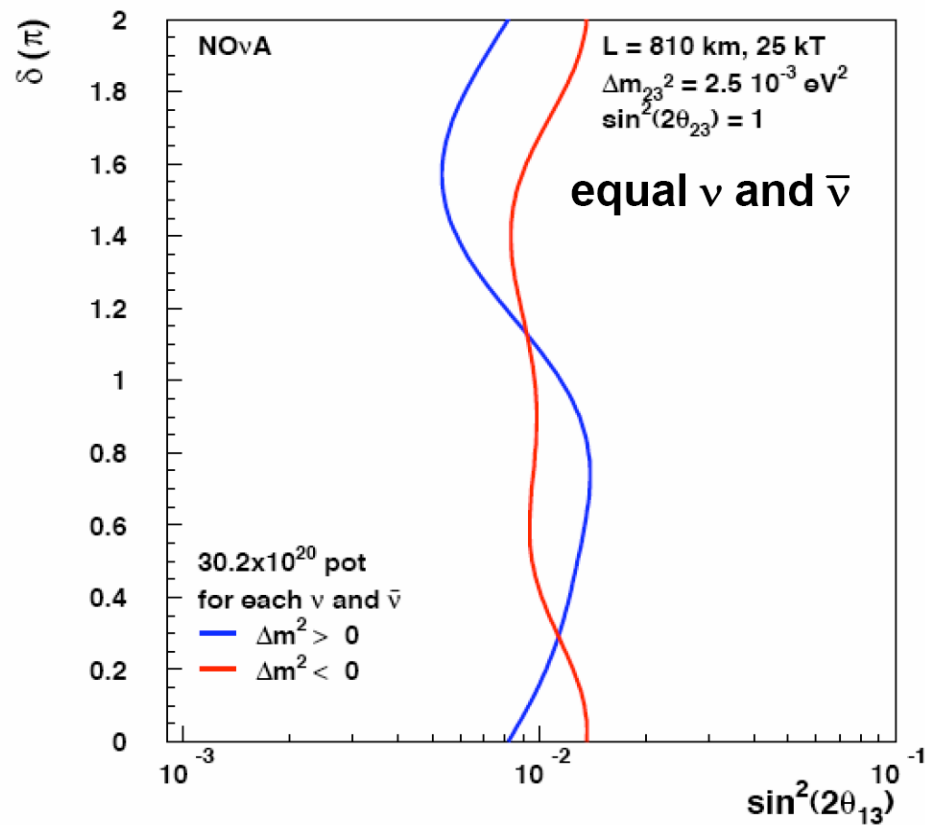
3 σ Sensitivity to $\theta_{13} \neq 0$



3 σ Sensitivity to $\sin^2(2\theta_{13}) \neq 0$



3 σ Sensitivity to $\sin^2(2\theta_{13}) \neq 0$

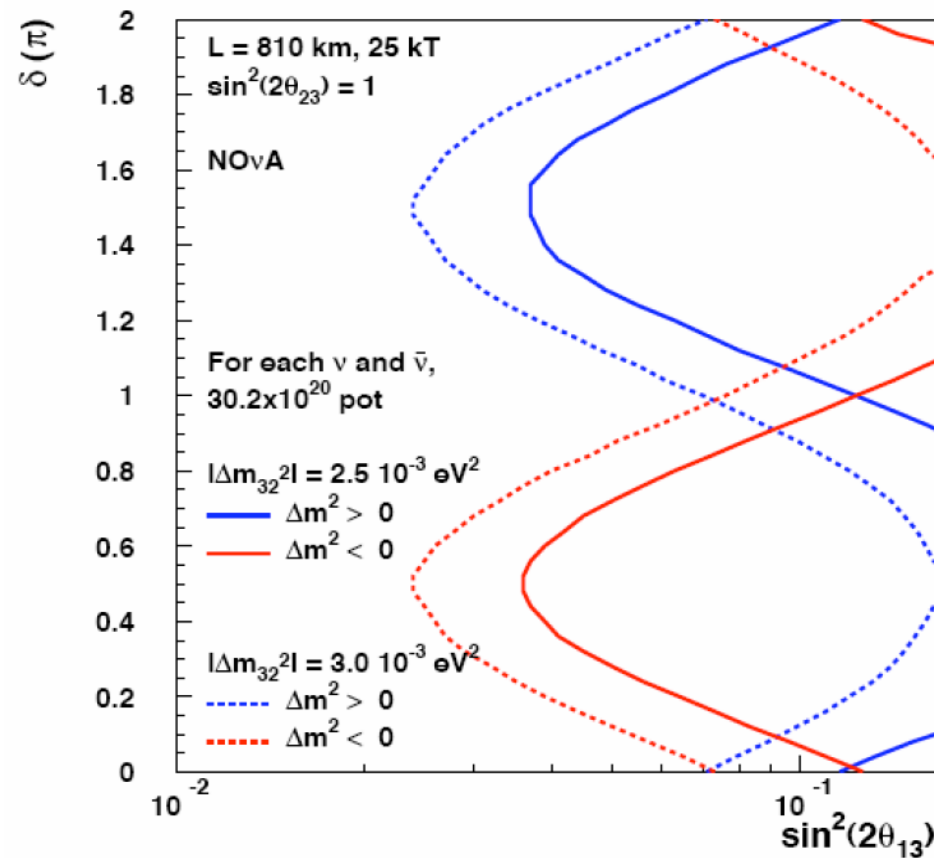




Resolution of the mass hierarchy



95% CL Resolution of the Mass Hierarchy





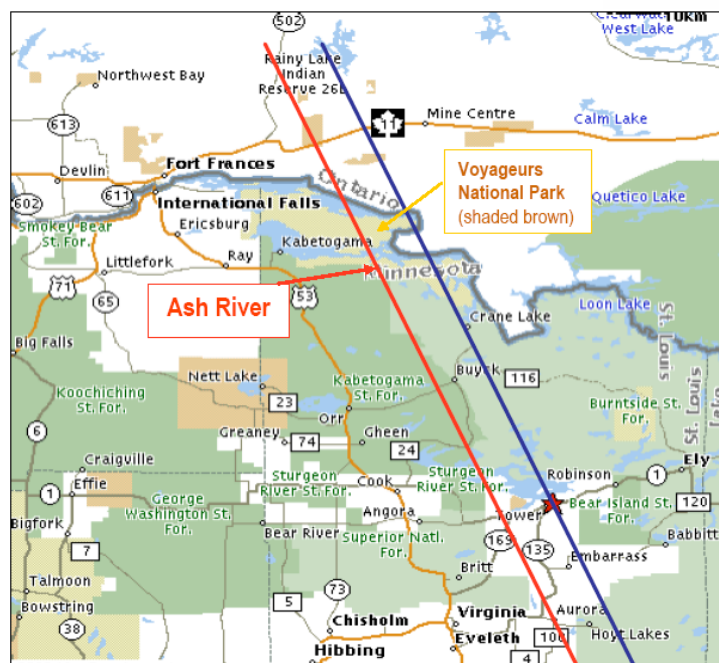
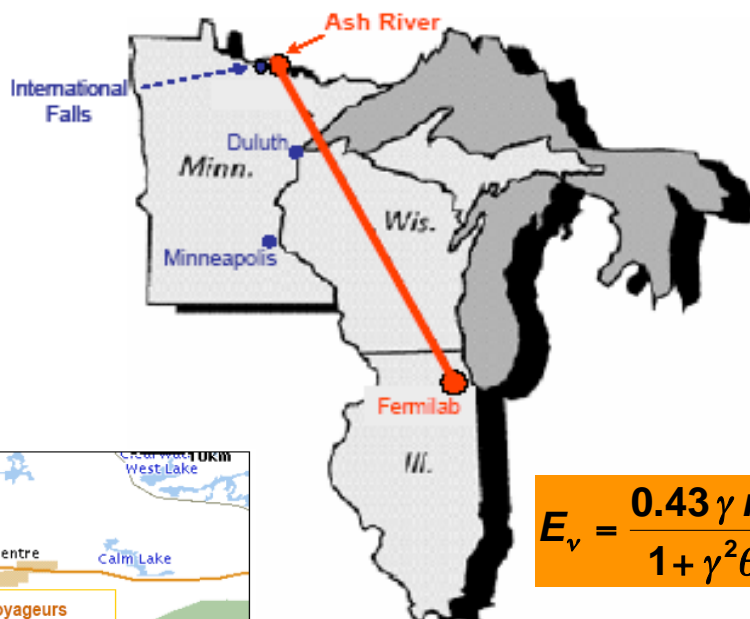
The strategy: off-axis NuMI beam



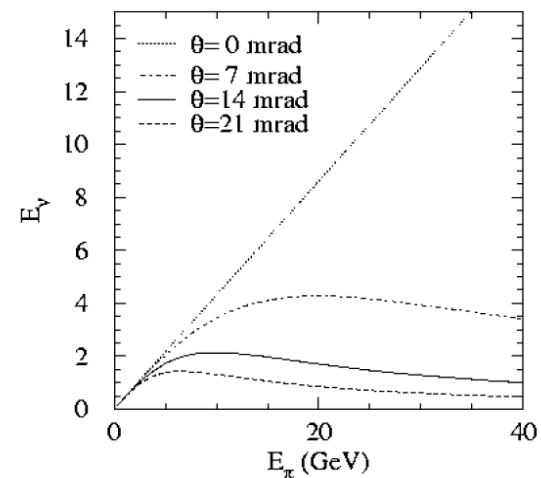
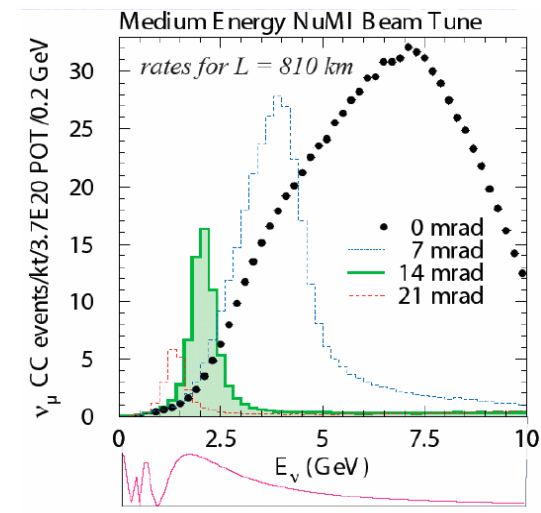
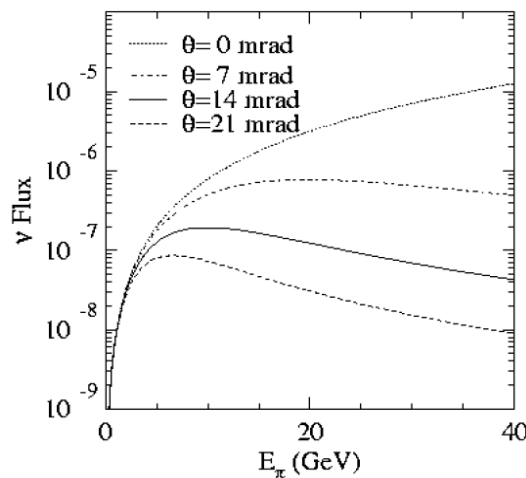
❑ Fermilab – Ash River

❑ 14 mrad off-axis

❑ 810 km baseline



$$E_\nu = \frac{0.43 \gamma m_\pi}{1 + \gamma^2 \theta^2}$$





Experimental setup: the Main Injector now



(Main Injector = MI)

- ❑ MI is fed 1.56 μ s batches from 8 GeV Booster

(MI ramp time ~ 1.5 sec)

- ❑ NuMI designed for

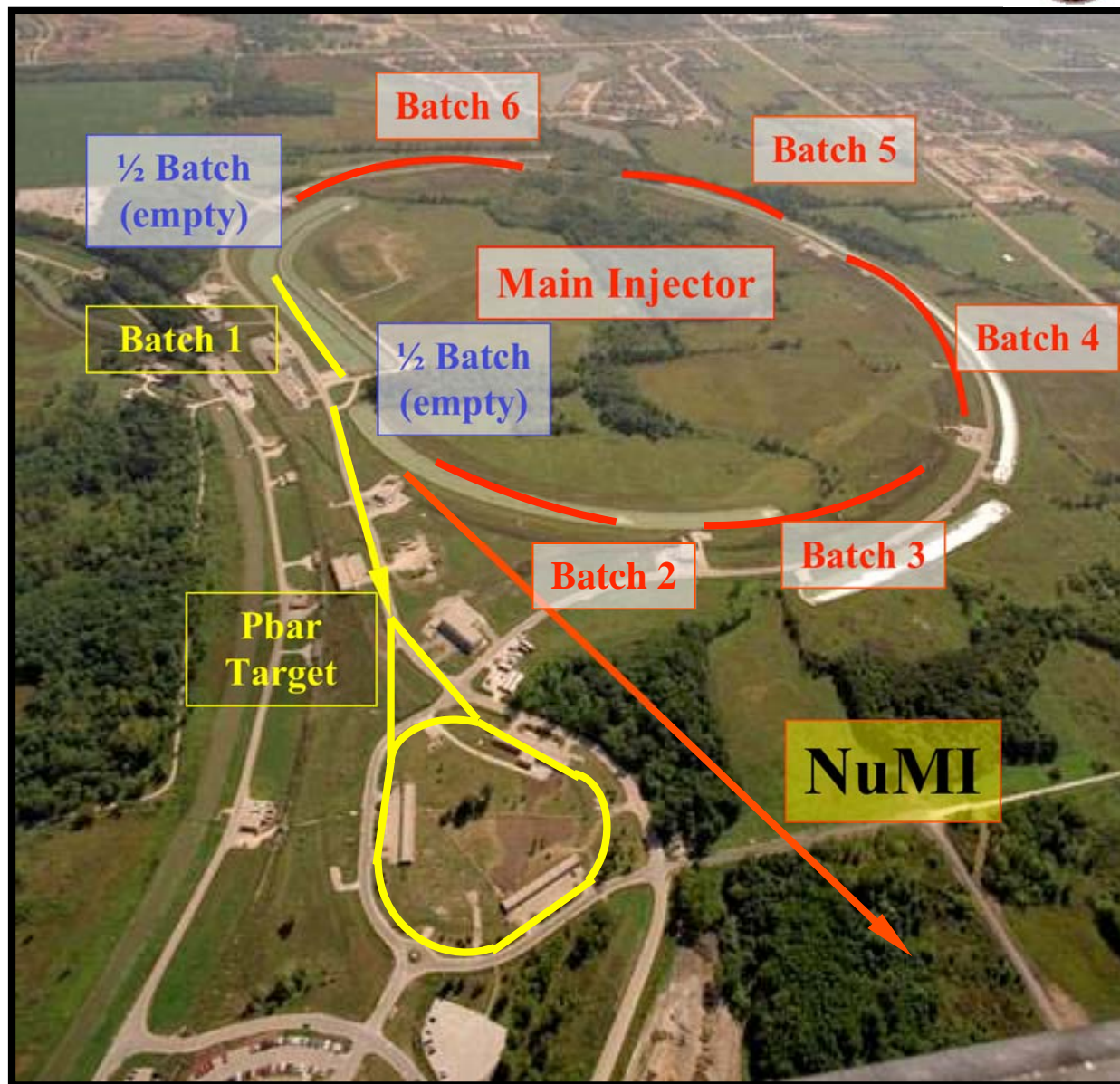
- ◆ 9.6 μ sec single turn extraction
- ◆ 4×10^{13} ppp @ 120 GeV
- ◆ 1.9 second cycle time
- ◆ beam power 400kW

- ❑ Typical performance to date:

- ◆ 2.4×10^{13} ppp @ 120 GeV
- ◆ 2.2-2.4 second cycle time

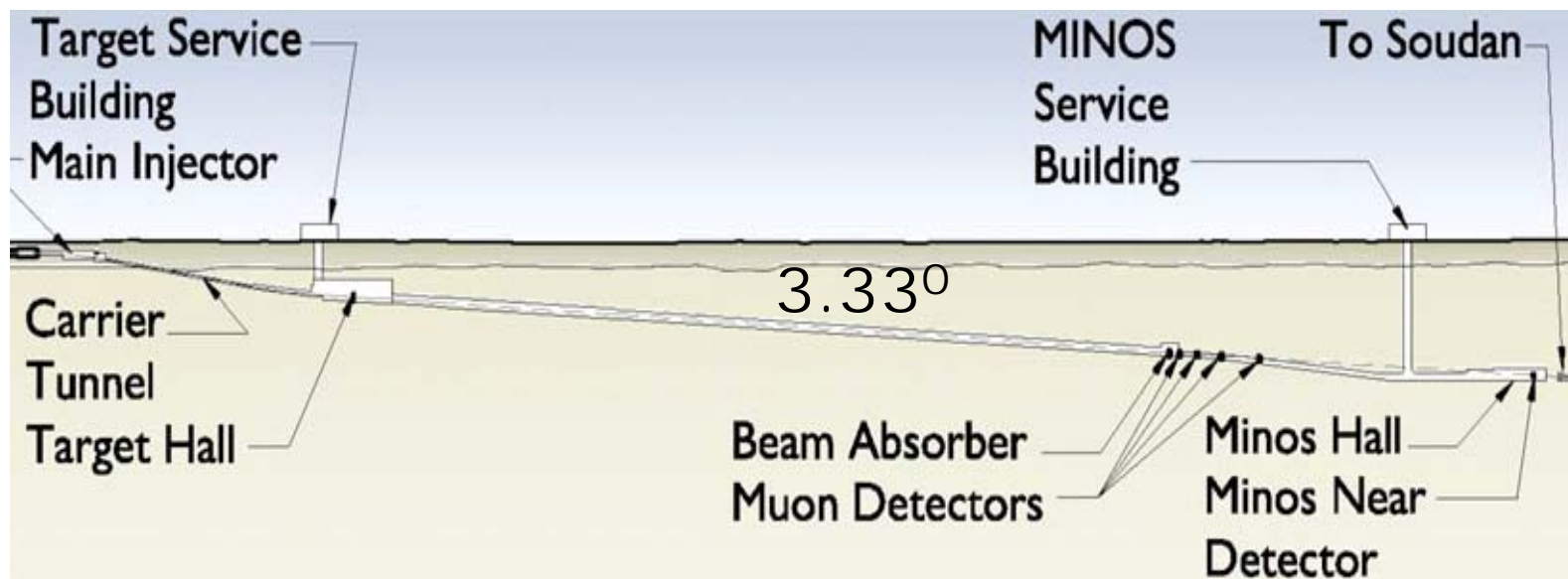
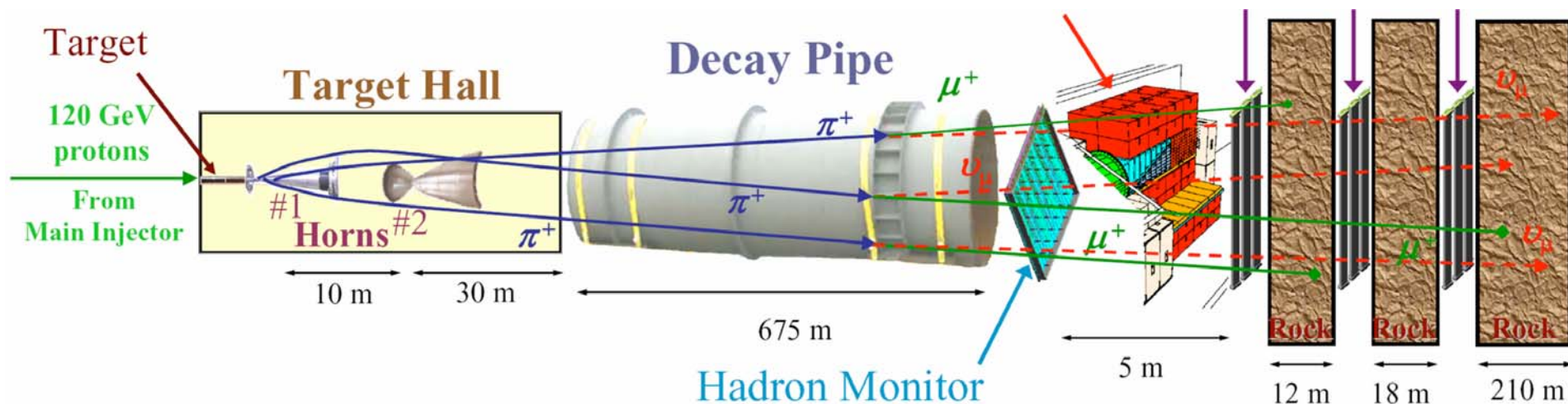
- ❑ Achieved records:

- ◆ 4.05×10^{13} ppp @ 120 GeV
(Feb 22, 2007)
- ◆ 2.0 second cycle time
- ◆ 315 kW



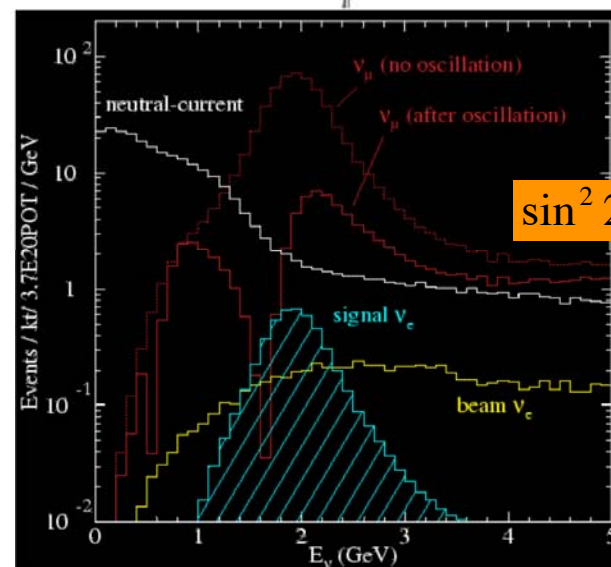
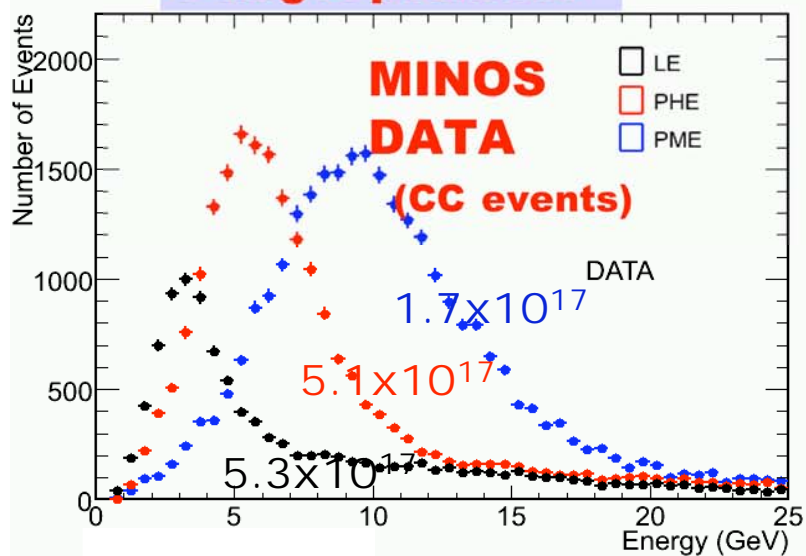
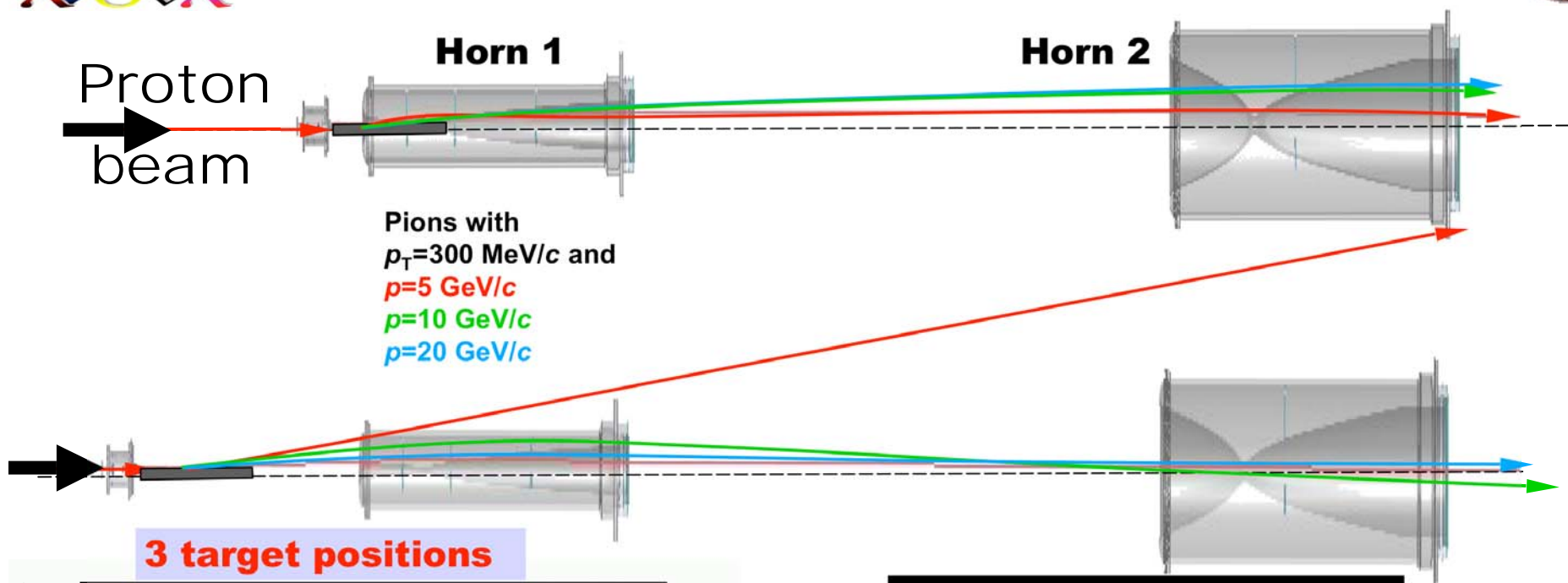


Experimental setup: NuMI beam





NuMI – multi-beam



$$\sin^2 2\theta_{13} = 0.04$$



Proton Improvement Plan



□ More protons with Collider (<2009)

- ◆ 9/11 slip-stacked Booster Batches (2 batches for anti-protons)
- ◆ Repetition rate = 0.8 s (Booster) + 1.4 s (MI ramp) = 2.2 s
- ◆ 3.4×10^{20} protons/year

□ Post-Collider era w/ Recycler (2010-2011)

(Accelerator and NUml Upgrade [ANU])

- ◆ 12 batches
- ◆ use Recycler for slip stacking (1.33 s cycle)
- ◆ 6×10^{20} protons/year

□ Post-Collider era w/ Accumulator (>2012)

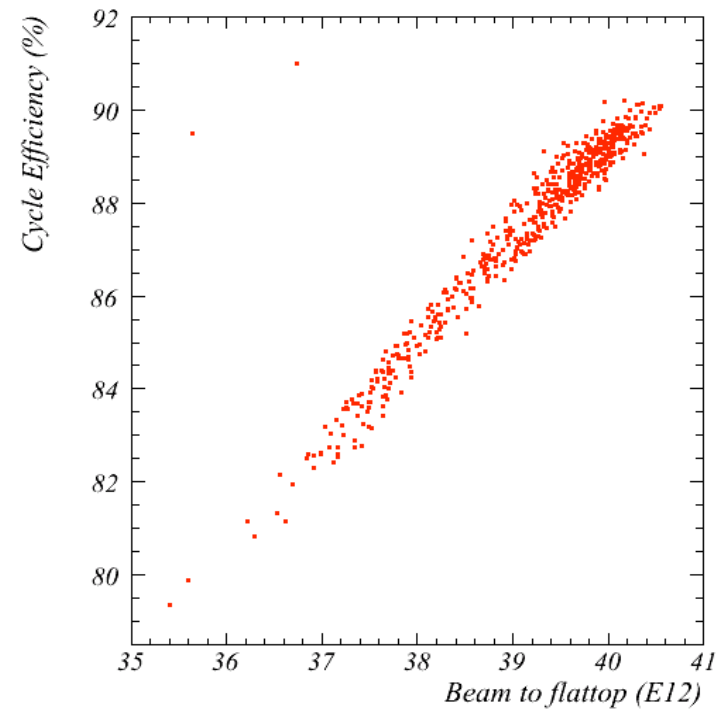
- ◆ Use Accumulator (for momentum stacking)
- ◆ 10×10^{20} protons/year



Very recent progress



❑ 4.05×10^{13} ppp slip-stacked in 11 batches (on Feb 22, 2007)





Operating Scenarios



	Present operating conditions *	Proton Plan Multi -batch slip-stacking in MI *	NOvA ANU Multi -batch slip -stacking in Recycler	Proposed SNuMI Accumulator momentum stacking
Booster intensity (p/batch)	$4.5 \cdot 10^{12}$	$4.3 \cdot 10^{12}$	$4.3 \cdot 10^{12}$	$4.7 \cdot 10^{12}$
No. Booster batche s to NuMI	5	9	12	18
MI cycle time (s)	2	2.2	1.333	1.333
MI intensity (ppp)	$3.1 \cdot 10^{13}$	$4.5 \cdot 10^{13}$	$4.9 \cdot 10^{13}$	$8.3 \cdot 10^{13}$
To NuMI (ppp)	$2.25 \cdot 10^{13}$	$3.7 \cdot 10^{13}$	$4.9 \cdot 10^{13}$	$8.3 \cdot 10^{13}$
NuMI beam power (kW)	210	320	700	1200
POT/yr to NuMI	$2 \cdot 10^{20}$	$3 \cdot 10^{20}$	$6 \cdot 10^{20}$	$10 \cdot 10^{20}$
MI protons/hr	$5.5 \cdot 10^{16}$	$7.3 \cdot 10^{16}$	$1.3 \cdot 10^{17}$	$2.2 \cdot 10^{17}$

* NuMI values are given for mixed-mode cycles

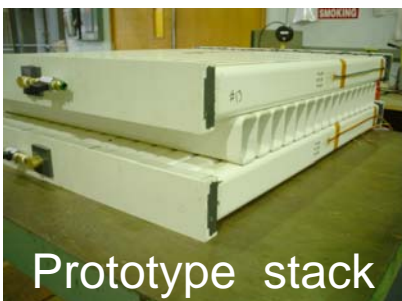


NOvA Far Detector we would like to build



We will build as much of this as the funding will allow...

- ❑ TAD = Totally Active Detector
PVC = passive material
- ❑ mass N kT (N large)
~80% scintillator
~20% PVC extrusions
- ❑ Modular structure
32 cells/extrusion
12 extrusions/plane
1984 planes
- ❑ Cell dimensions:
3.9 cm x 6 cm x 15.7m
- ❑ U-shaped 0.7 mm WLS fiber into APD
- ❑ $X_0 = 44$ cm $\rho_M = 10$ cm

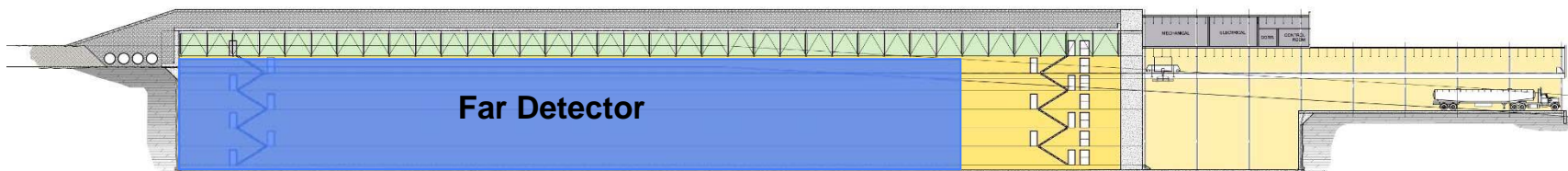
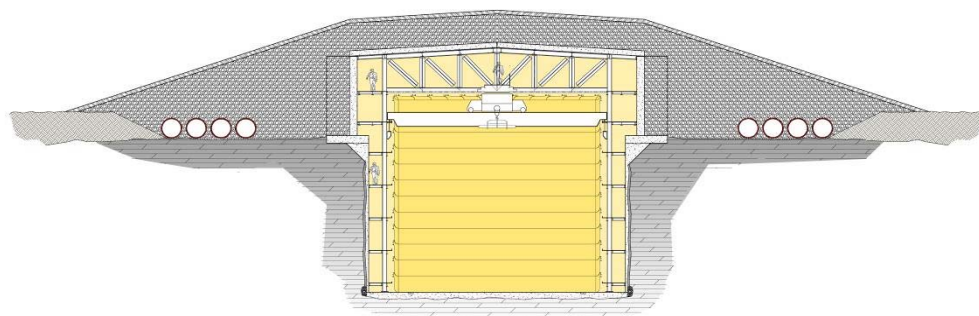


Prototype stack

15.7m

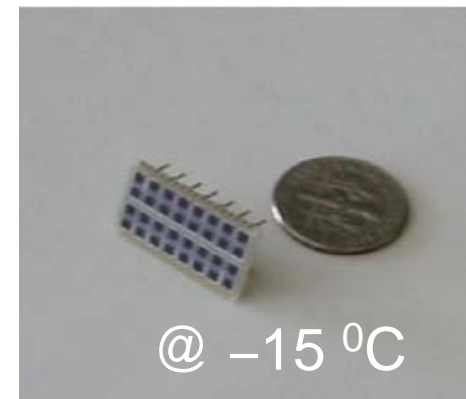
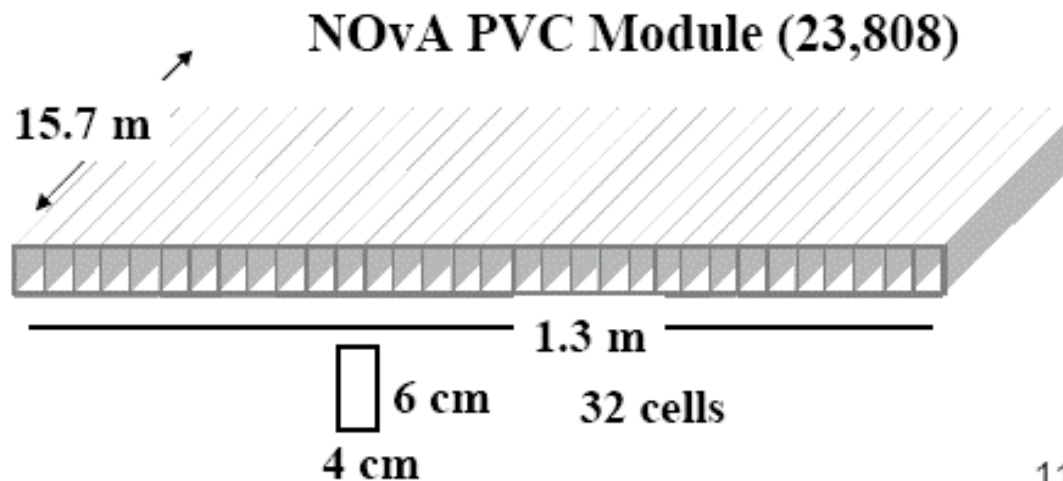
15.7m

31-plane
blocks
(4 blocks in a
"super-block")





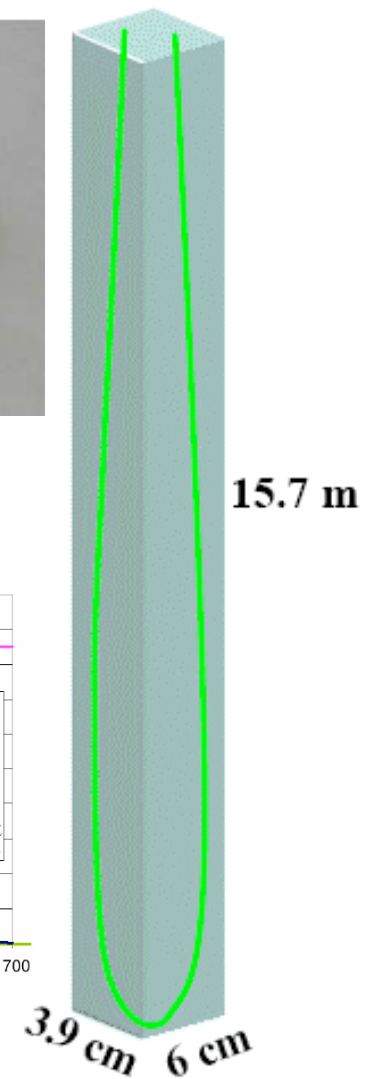
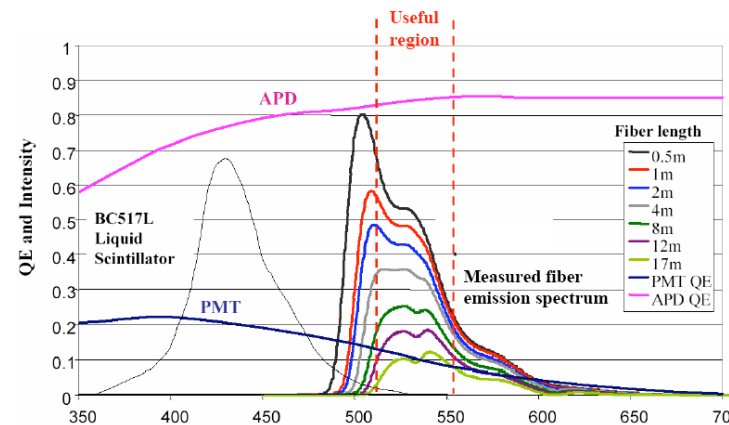
Detector technology



To 1 APD pixel

Active readout components:

- ❑ Liquid scintillator
 - filled cells.
- ❑ WLS fiber
 - 0.7 mm diameter
 - looped end ("perfect" reflector)
 - readout both ends on one side
- ❑ Avalanche Photodiode
 - Hamamatsu multi-pixel
 - 85% QE





Scintillator, light yield

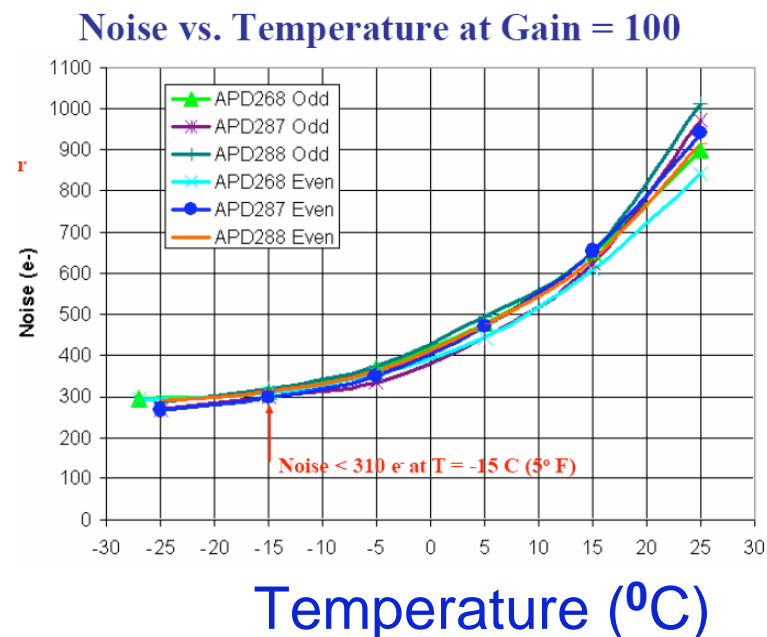
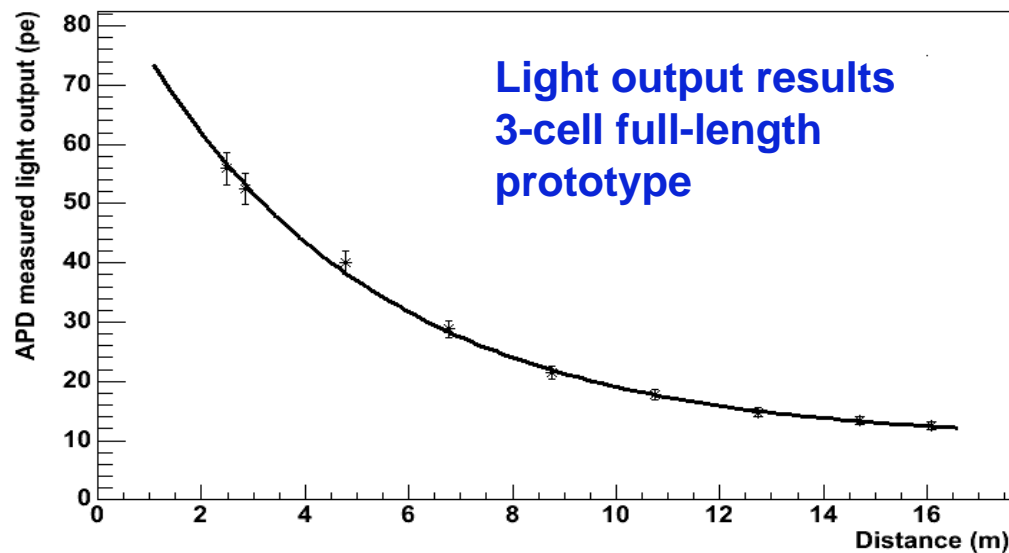


❑ NO_vA recipe

Equivalent to Saint-Gobain (Bicron) BC-517P or Eljen Technology EJ-321P

❑ Requirement: 20PE's for a MIP at far extrusion-end

Component	Purpose	Mass fraction
mineral oil	solvent	94.4%
pseudocumene	scintillant	5.5%
PPO	waveshifter #1	0.1%
bis-MSB	waveshifter #2	0.002%
Stadis-425	anti-static agent	0.0003%
tocopherol	anti-oxidant	0.0010%
TOTAL		100%



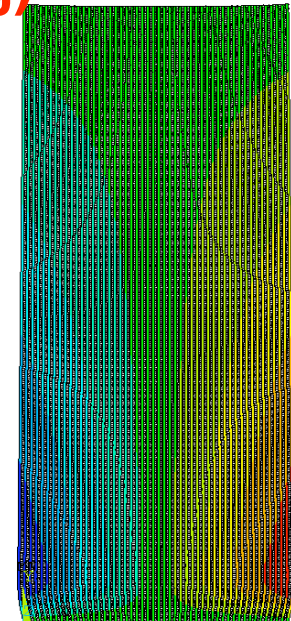
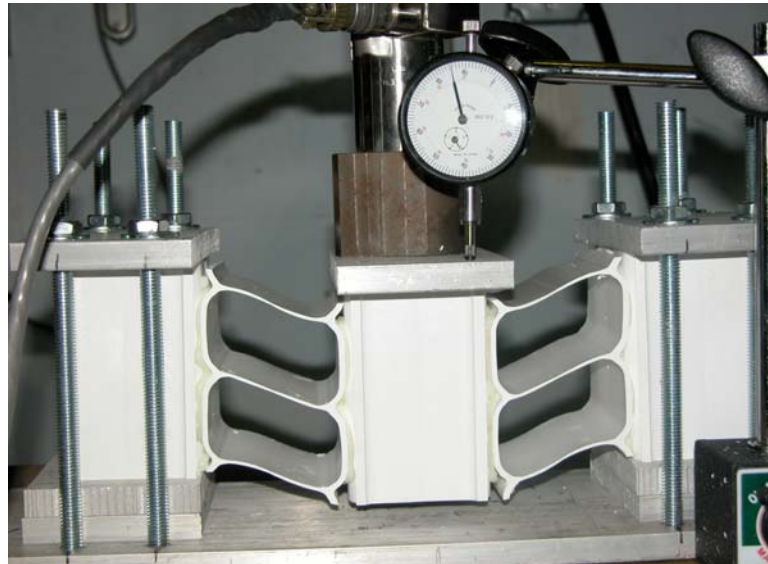


Structural challenges



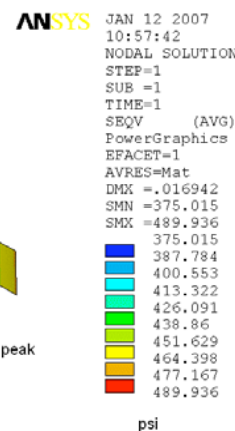
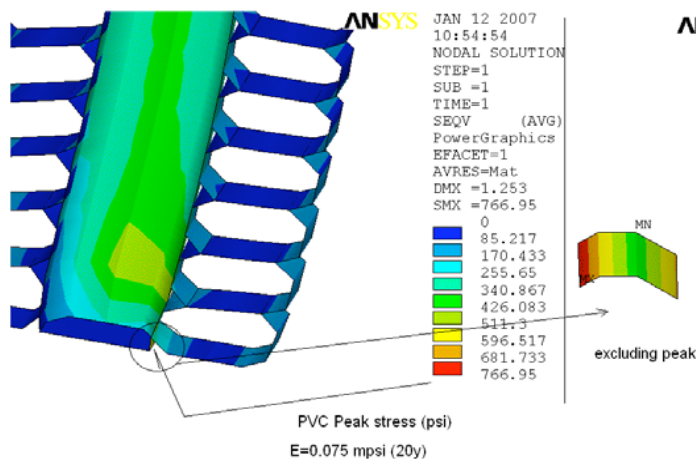
□ FEA calculations and tests (on-going)

31-plane block



```
ANSYS 11.0BETA
OCT 30 2006
11:00:25
NODAL SOLUTION
STEP=1
SUB =1
TIME=1
UX          (AVG)
RSYS=0
PowerGraphics
EFACET=1
AVRES=Mat
DMX =.47318
SMN =-.079761
SMX =.079761
-.079761
-.062036
-.044312
-.026587
-.008862
.008862
.026587
.044312
.062036
.079761
inch
```

deflection for 99 planes
block

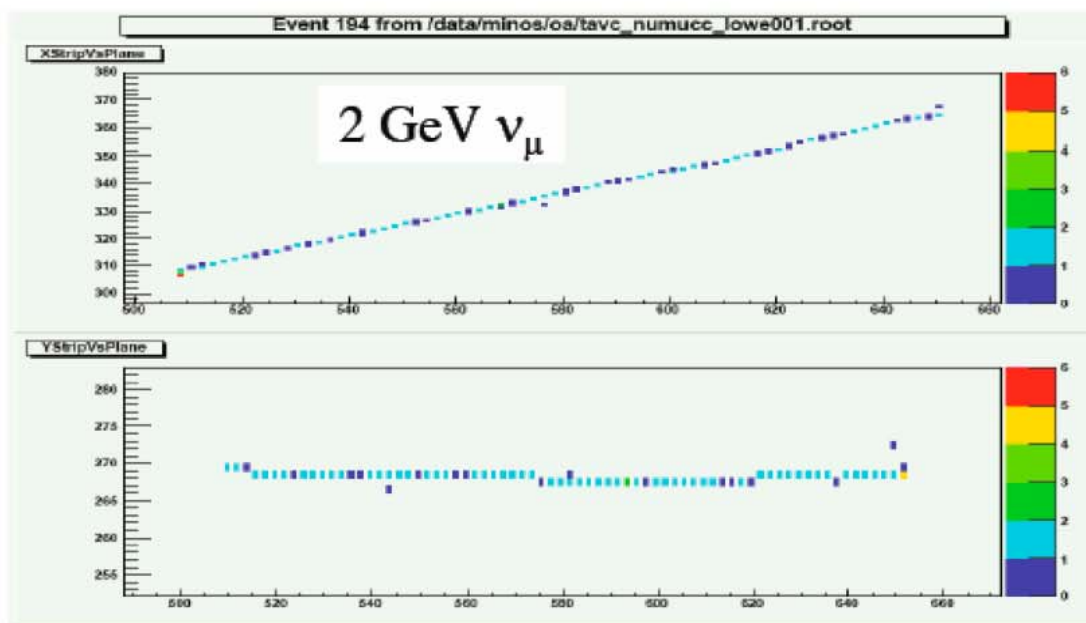
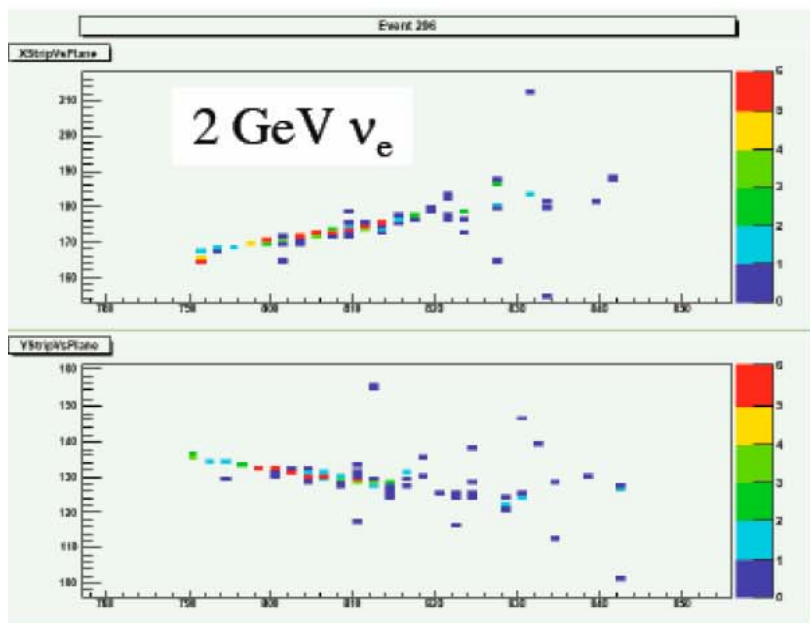




Event classification

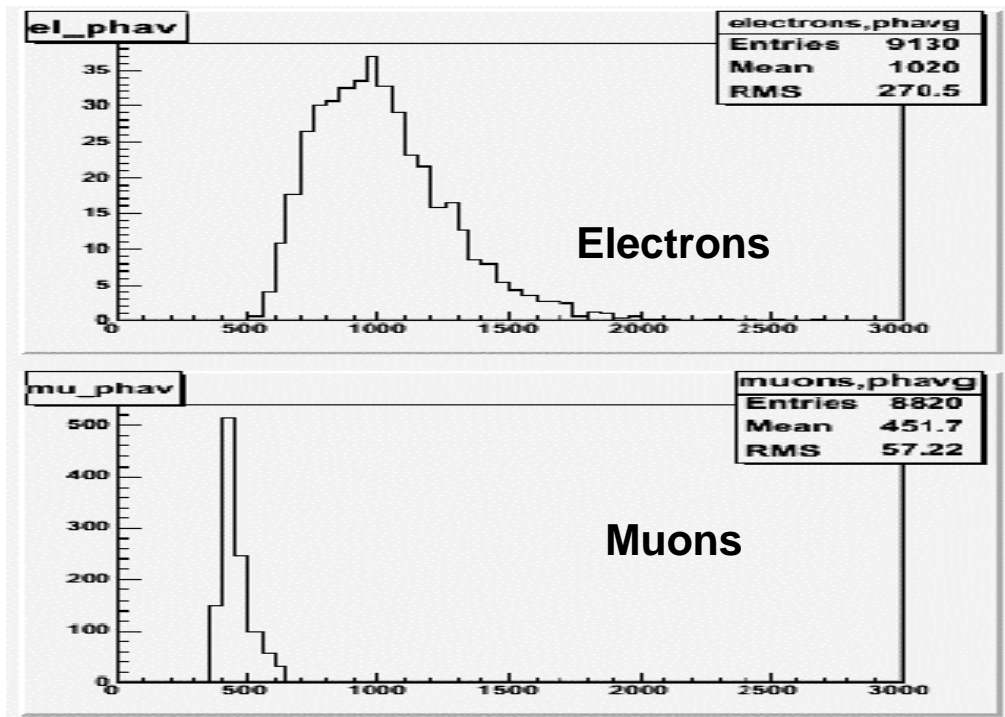


- Longitudinal sampling every 0.15 X0
- 2 GeV muon traverses ~60 planes

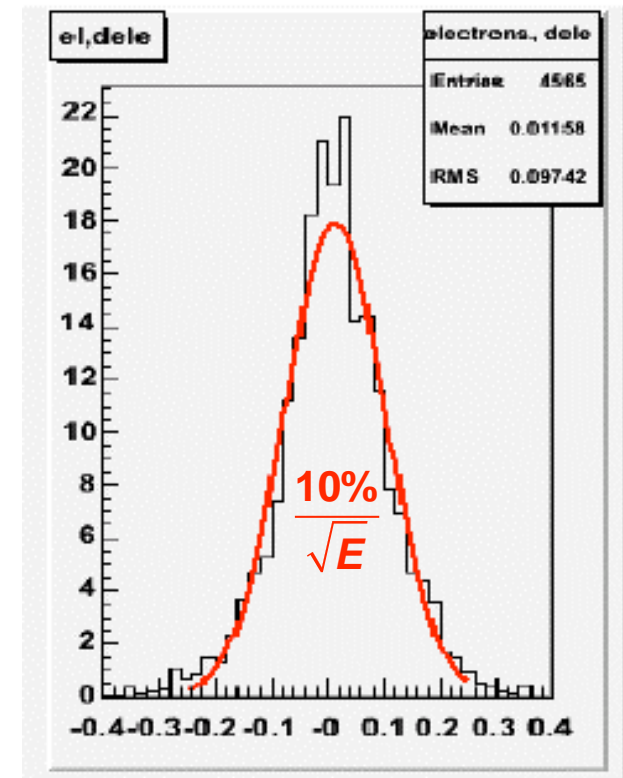




Electron ID and Resolution



Average pulse height per plane



Electron
energy
resolution



Near Detector

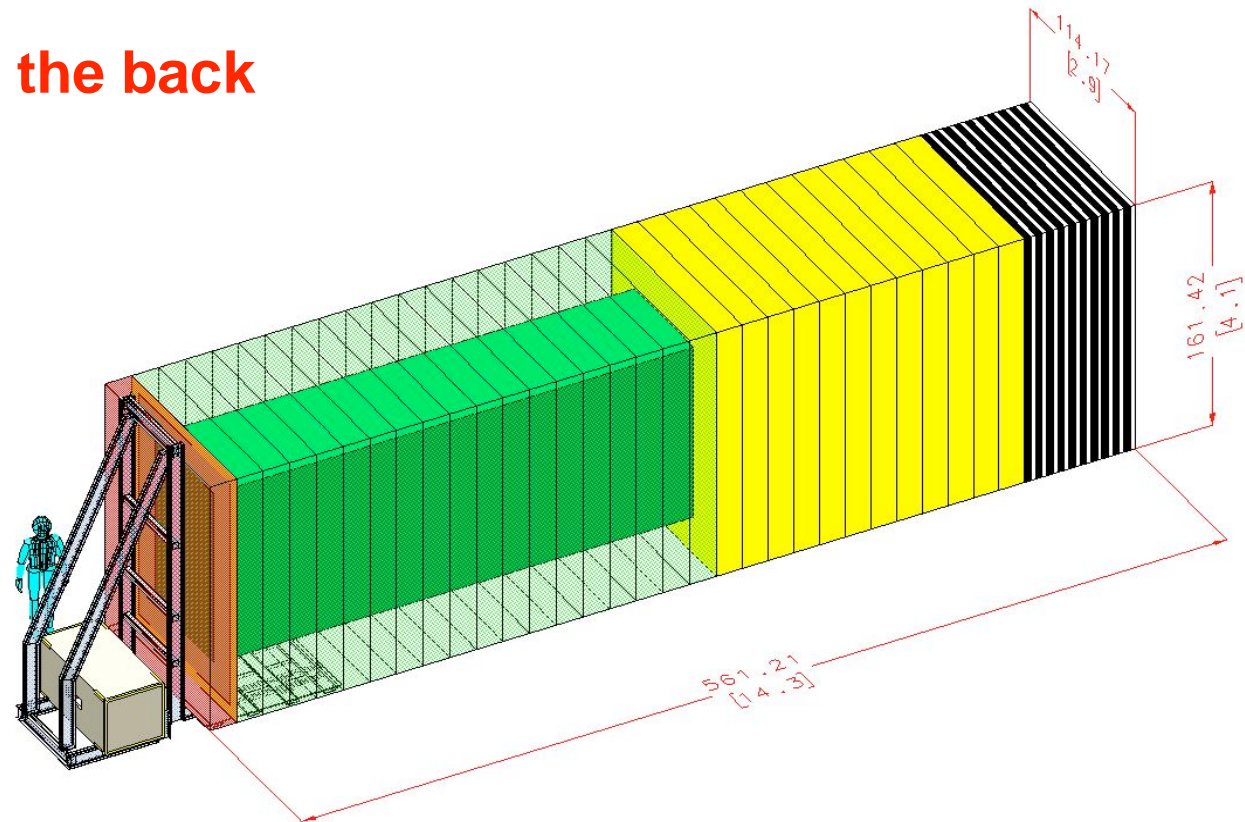


□ 2000 ν_e CC events per year

□ 20 tons fiducial volume

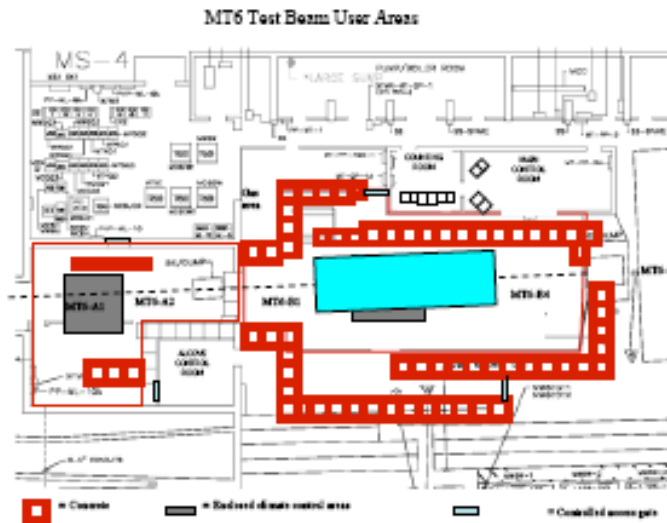
◆ 1.65 x 2.85 x 7.4 m³

□ Muon catcher in the back

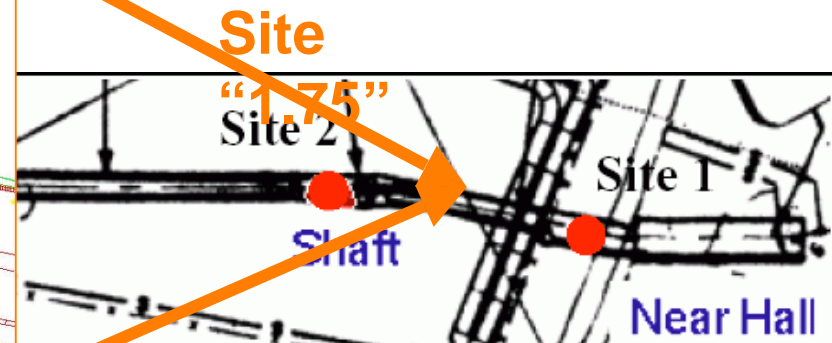
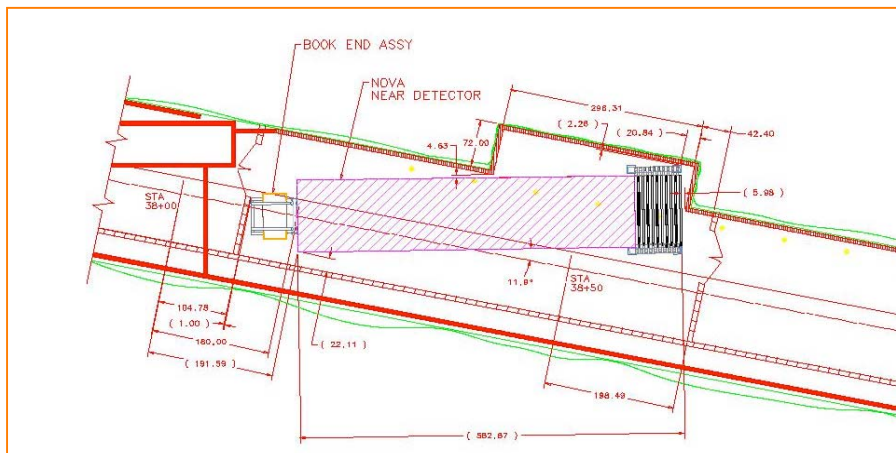
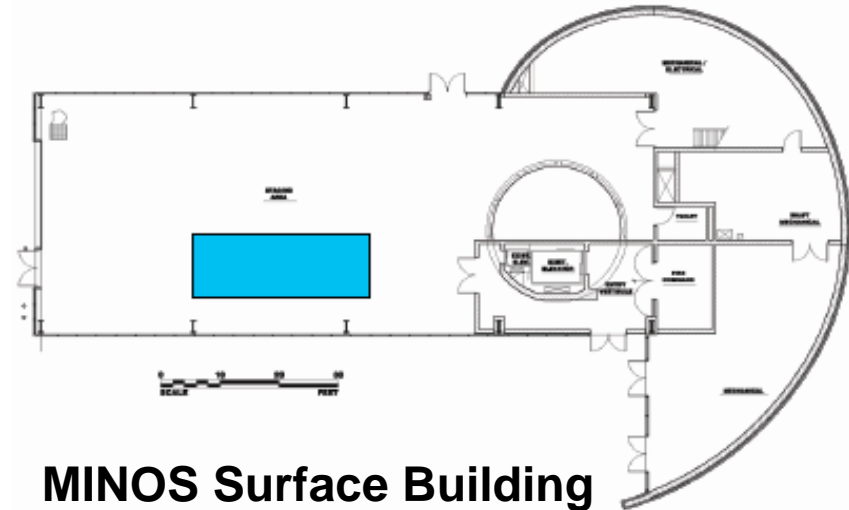




Near Detector in the beam(s)



M Test



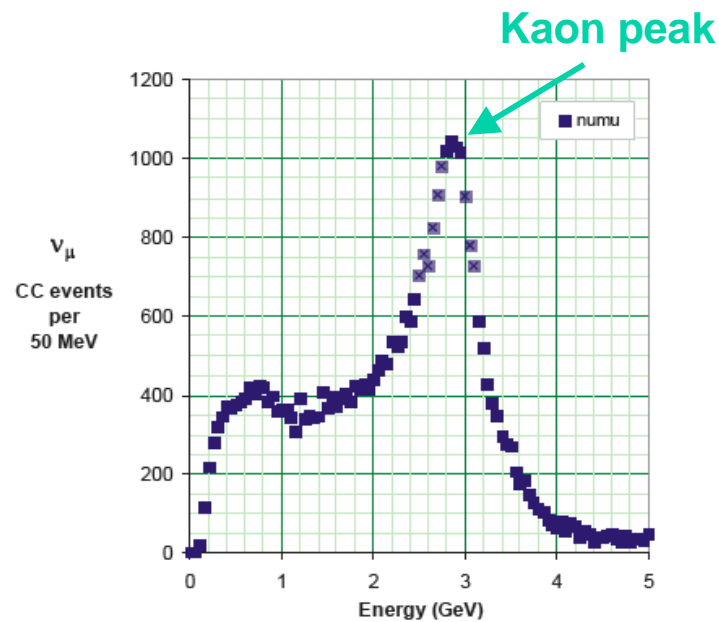
**NuMI Access Tunnel
(100 m underground)**



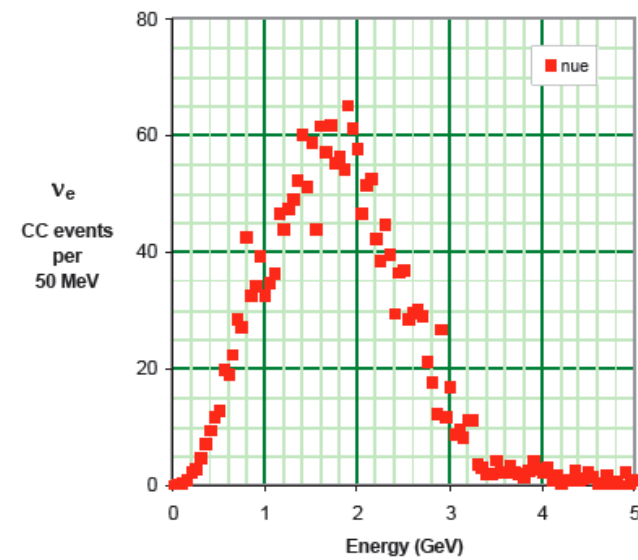
Near Detector in MINOS Surface Building



6.5×10^{20} pot in 75 mrad off-axis beam



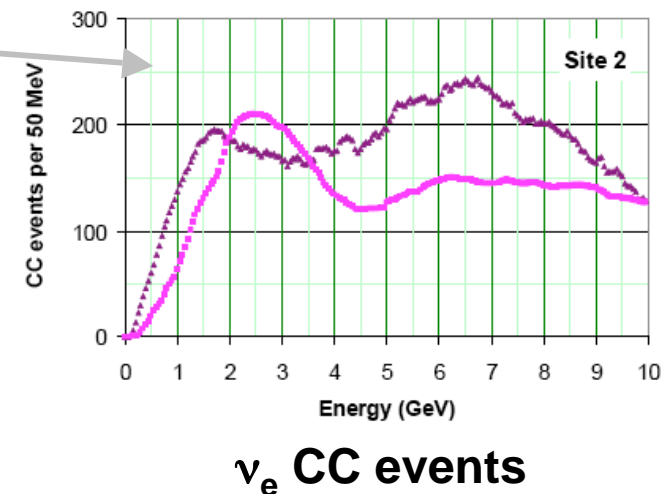
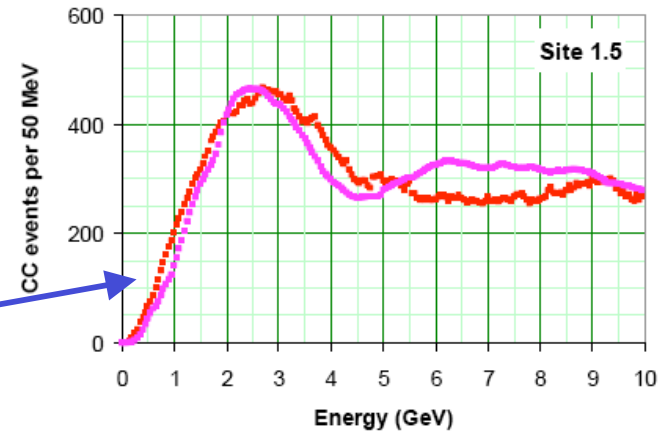
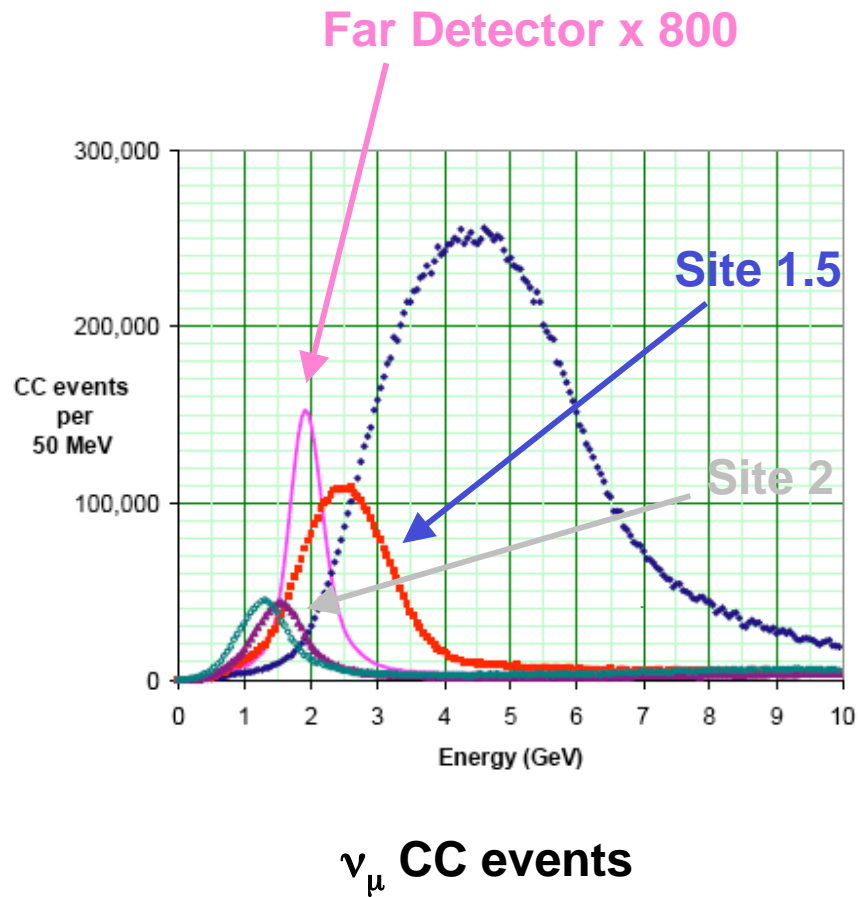
45,000 ν_μ CC events



2,200 ν_e CC events



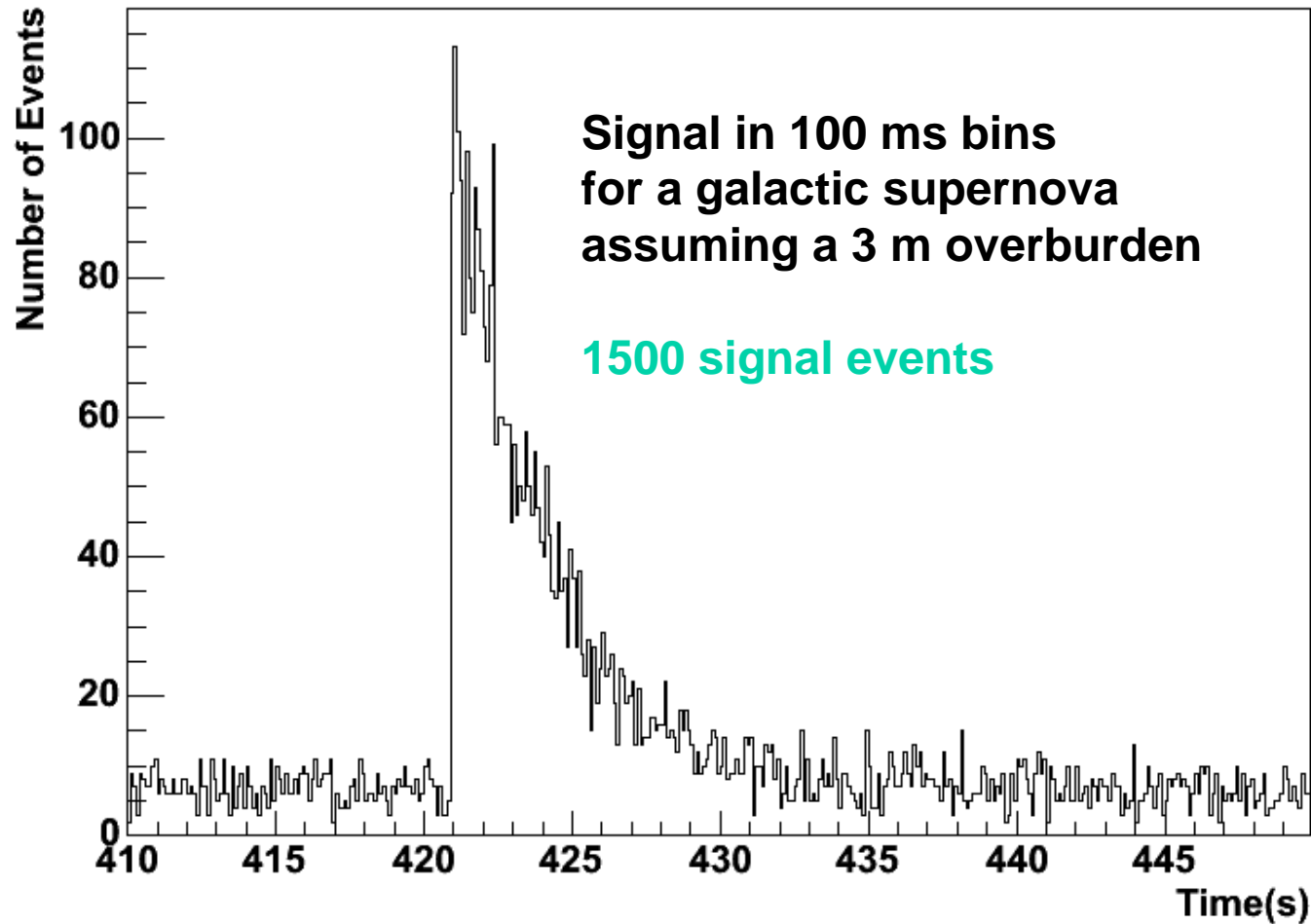
Near Detector in the NuMI Access Tunnel



For $6.5 \cdot 10^{20}$ pot



Sensitivity to a Galactic Supernova





Summary and outlook



□ NOvA will address central neutrino physics issues

- ◆ Subdominant mixing $\sin^2(2\theta_{13})$ to about 1-2%
- ◆ Neutrino mass hierarchy
- ◆ CP violation (δ in PMNS matrix)
- ◆ measure $\sin^2(2\theta_{32})$ to about 1-2%

□ Robust and straightforward detector design

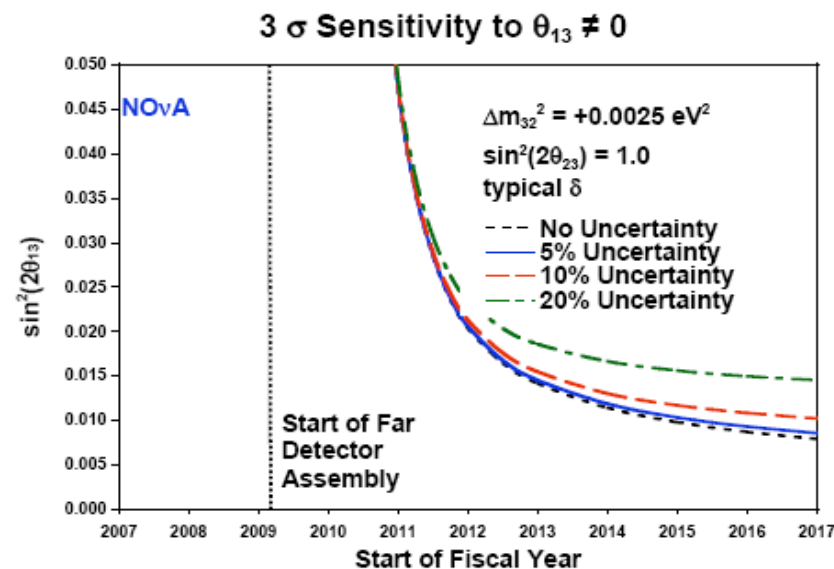
- ◆ In-house scintillator
- ◆ Long WLS fibers readout by APD's
- ◆ Modular (early start of data collection)

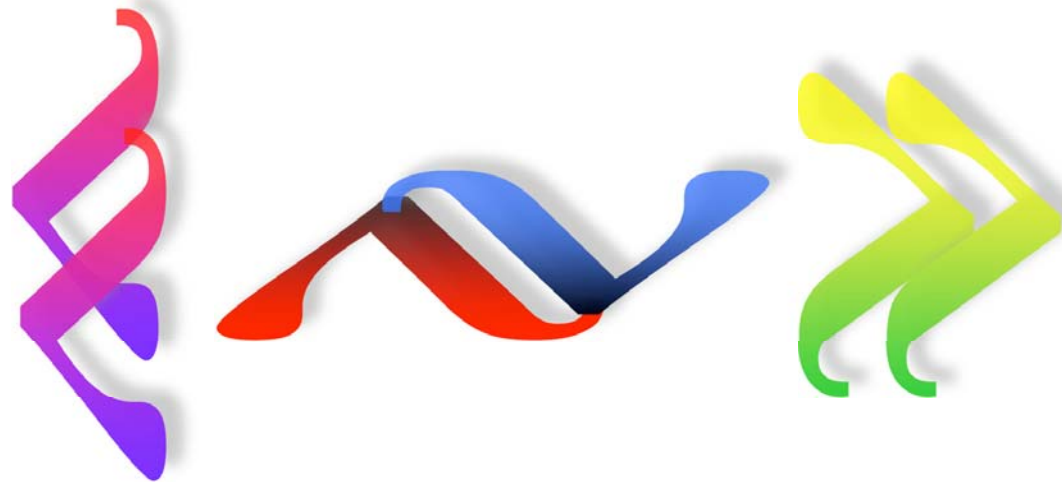
□ Accelerator & NuMI Upgrade

- ◆ 700 kW by 2009
- ◆ 1200 kW after 2012

□ Near term

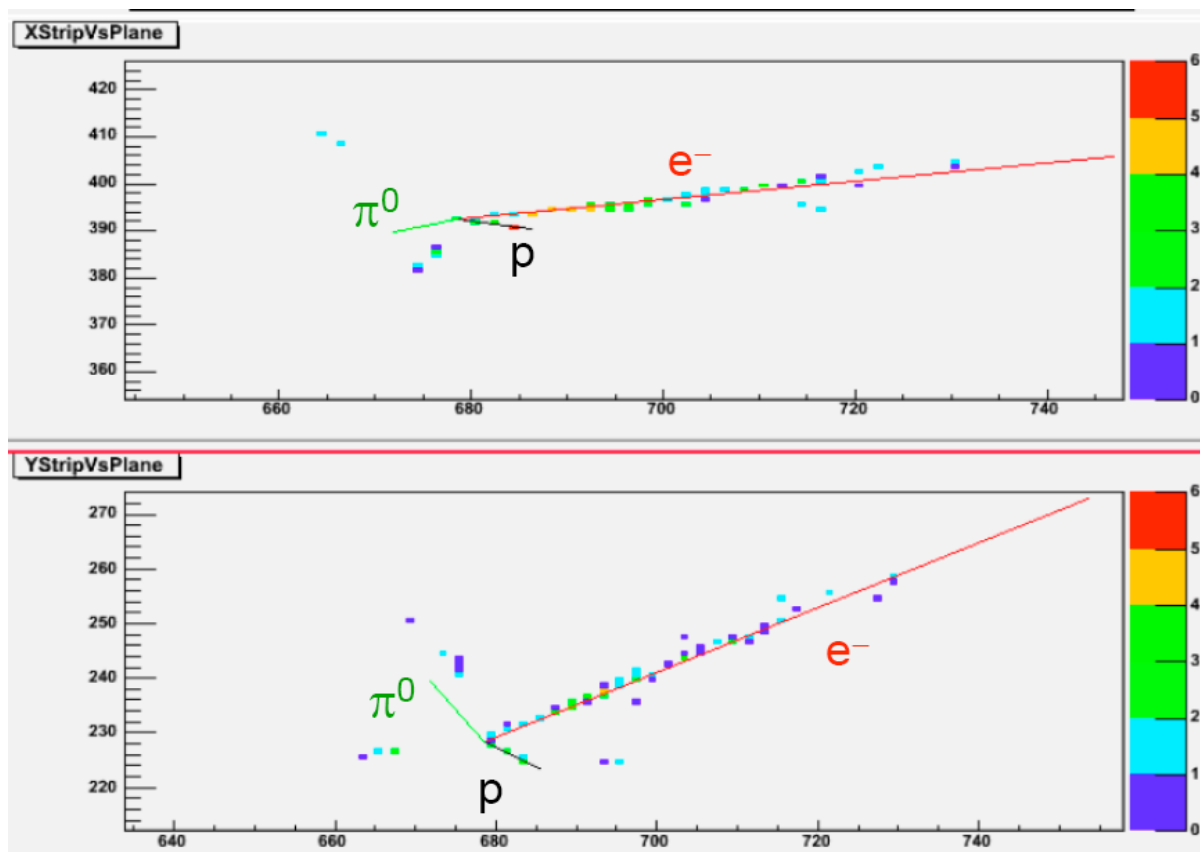
- ◆ NOvA in FY08 budget @ \$36M
- ◆ TDR
- ◆ prototype detector





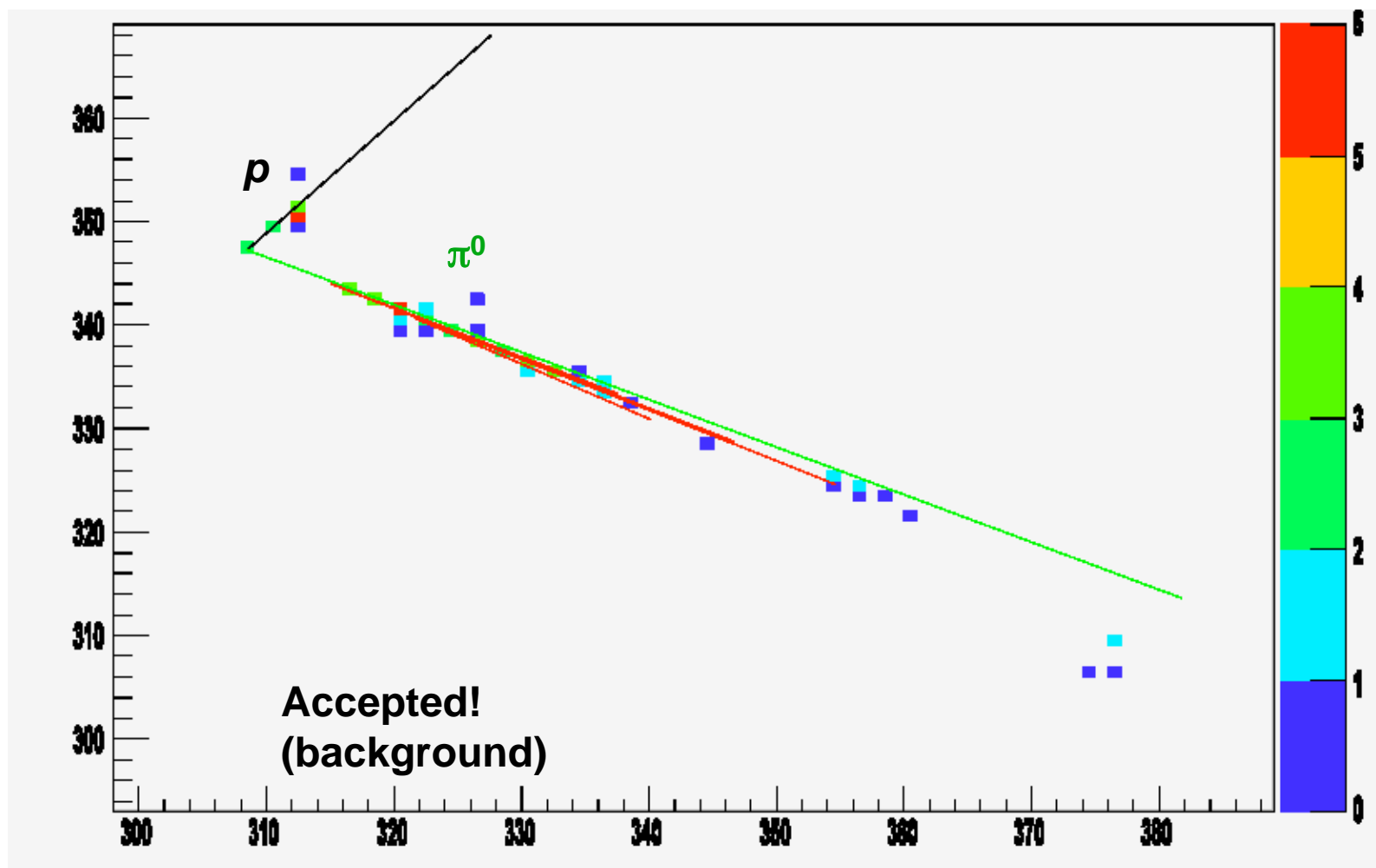


$1.65 \text{ GeV } \nu_e N \rightarrow e p \pi^0$



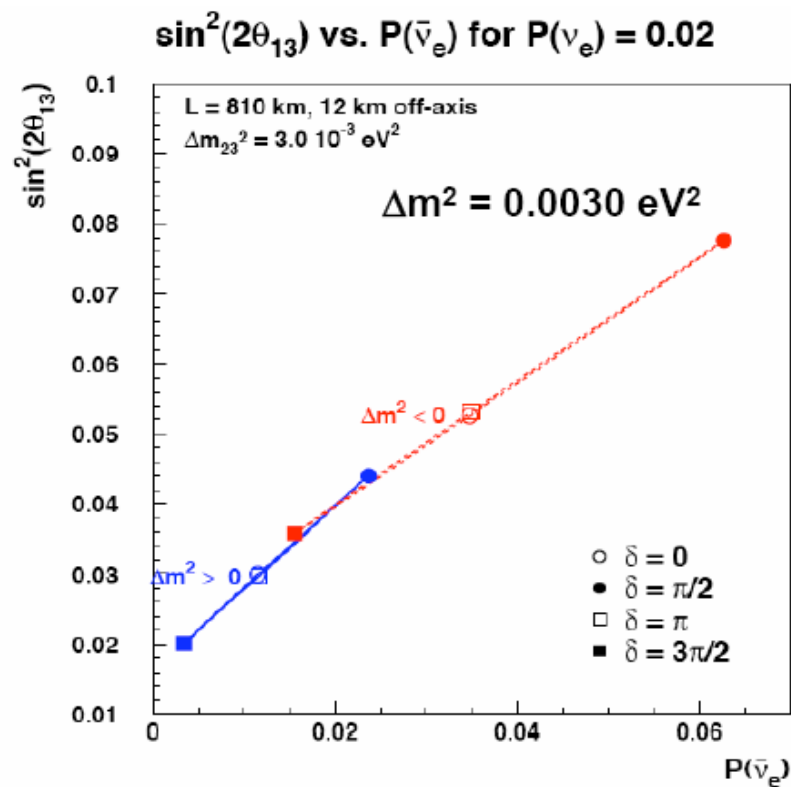
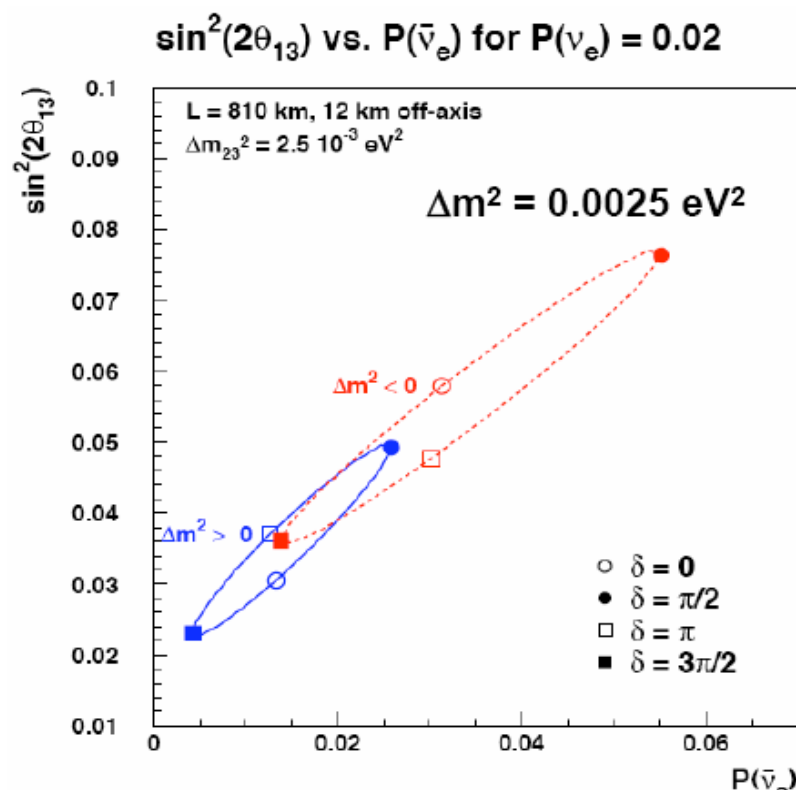


2.11 GeV $\nu_\mu N \rightarrow \nu_\mu p \pi^0$ x-z View



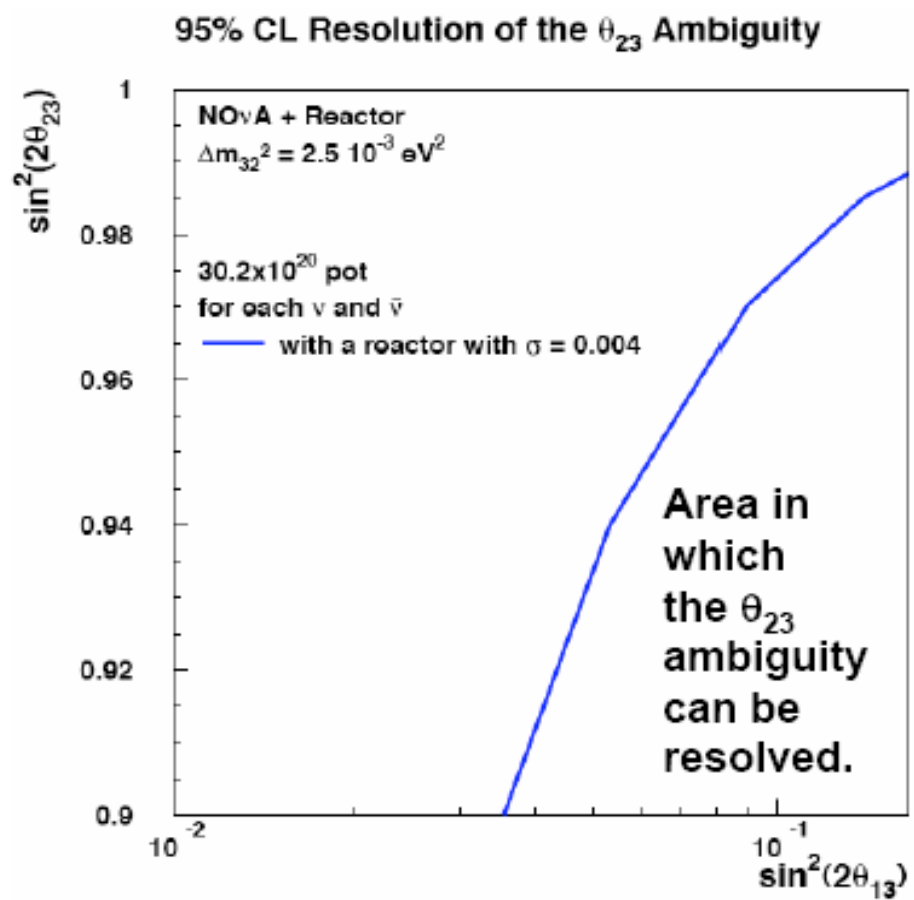


Ambiguities versus Δm_{23}^2





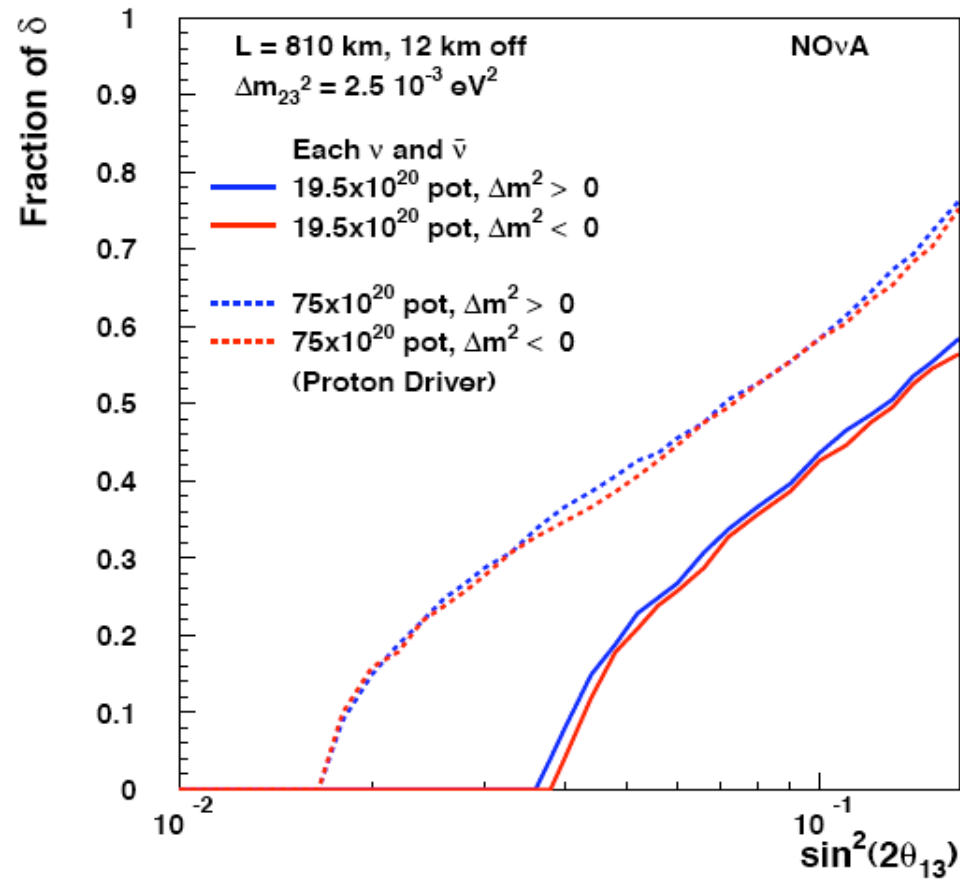
Resolution of the θ_{23} ambiguity



(There is some sensitivity to the mass ordering and δ . The blue line represents an average over these parameters.)

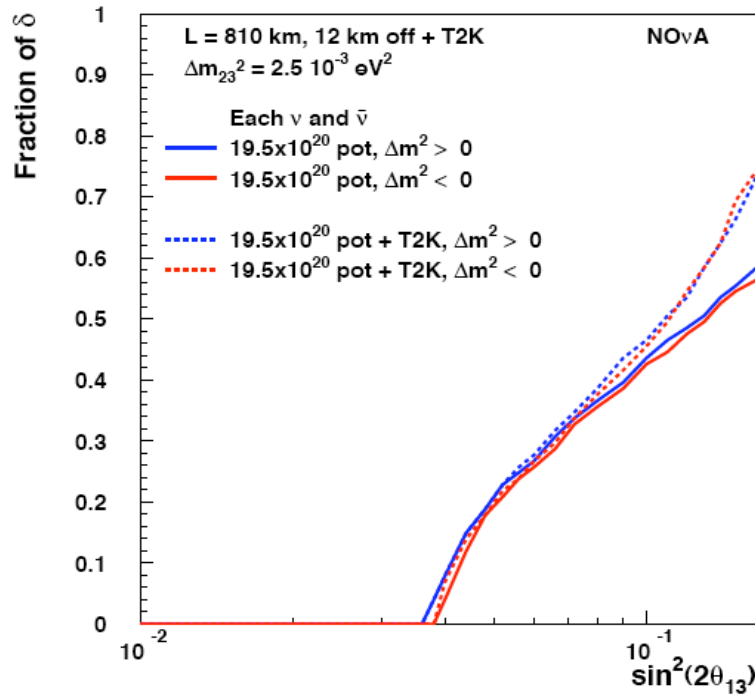


δ reach

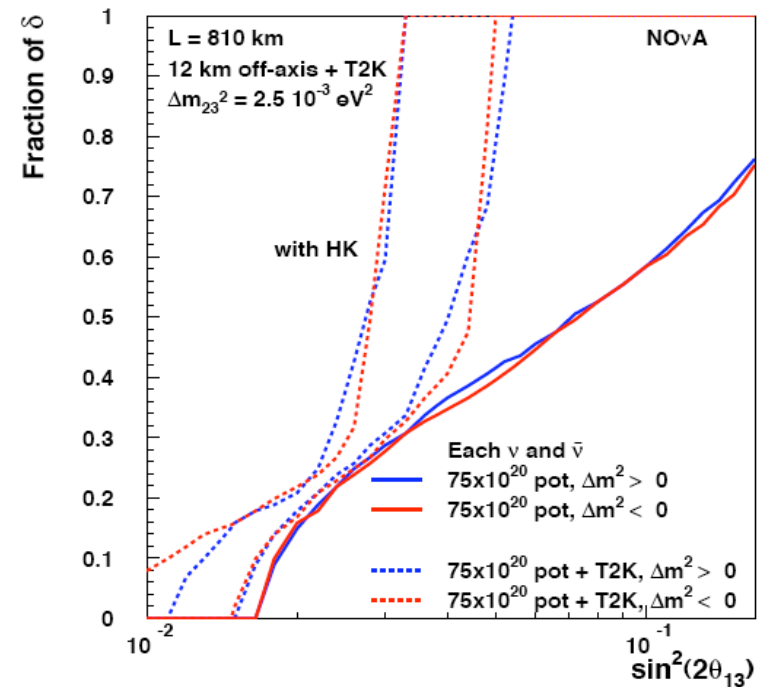




NOvA + T2K (to be updated)



NOvA with T2K Phase 1



NOvA/PD with T2K Phase 2

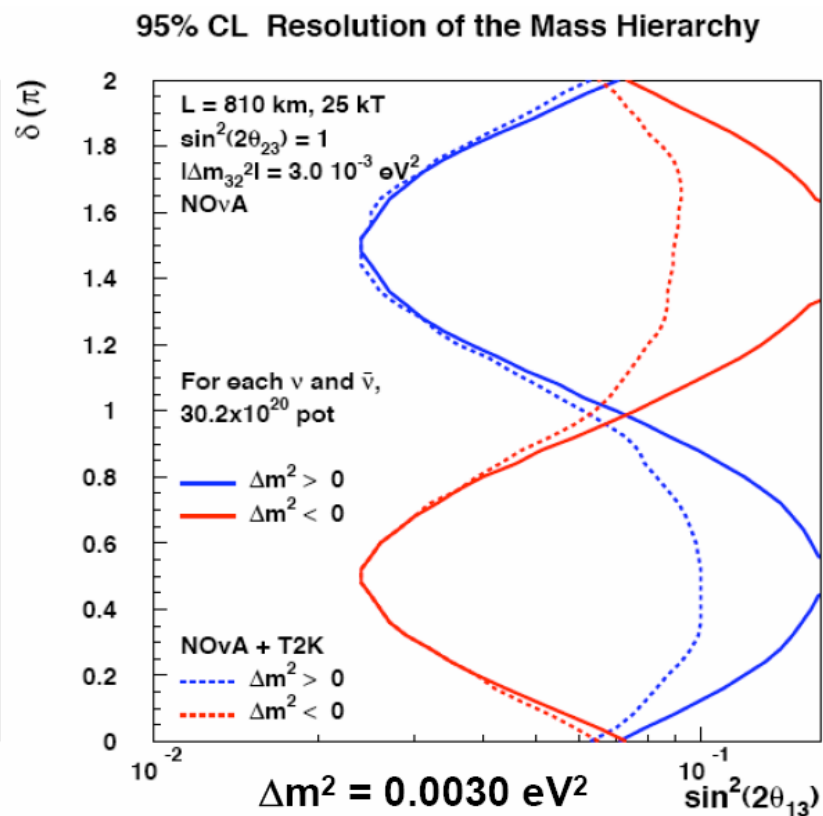
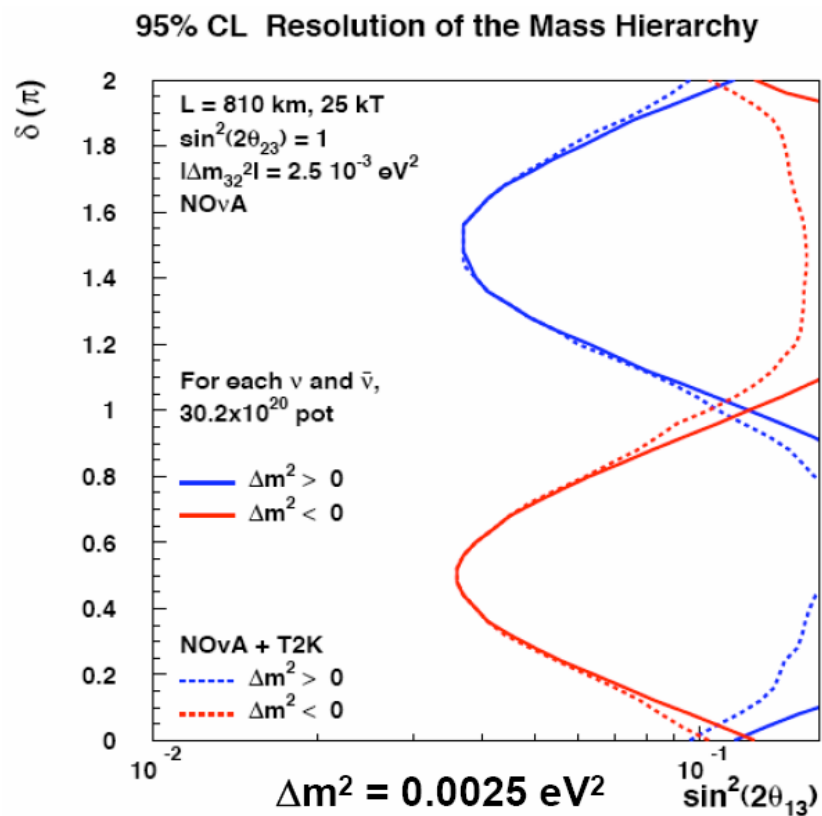
**95% CL Resolution of
the Mass Ordering**



Resolution of the mass hierarchy (with T2K)



□ To be updated





NOvA ANU possible Timeline



- ❑ 2008 shutdown: GPP Civil Work started for penetrations and overall GPP work completed in January 2009
- ❑ Fall 2009 TeV shutdown starts (assume)
 - ◆ Most likely accelerator installation not ready to start until Spring 2010 (low level of funding in FY08 delays completion of ANU work required before this shutdown)
 - ◆ To complete accelerator upgrades, need ~6-8 month shutdown
- ❑ Fall 2010: Start up and commission recycler with protons (400kW)
- ❑ 1 year of running with NuMI LE target for MINERvA
- ❑ Fall 2011 shutdown: complete NuMI Target Hall work for 700kW power and change to ME neutrino beam configuration
- ❑ January 2012: Start up and commission recycler (700kW)
- ❑ During 2011: NOvA 5kT detector ready